

# 長基線ニュートリノ振動実験 (K2K) 実験の現状

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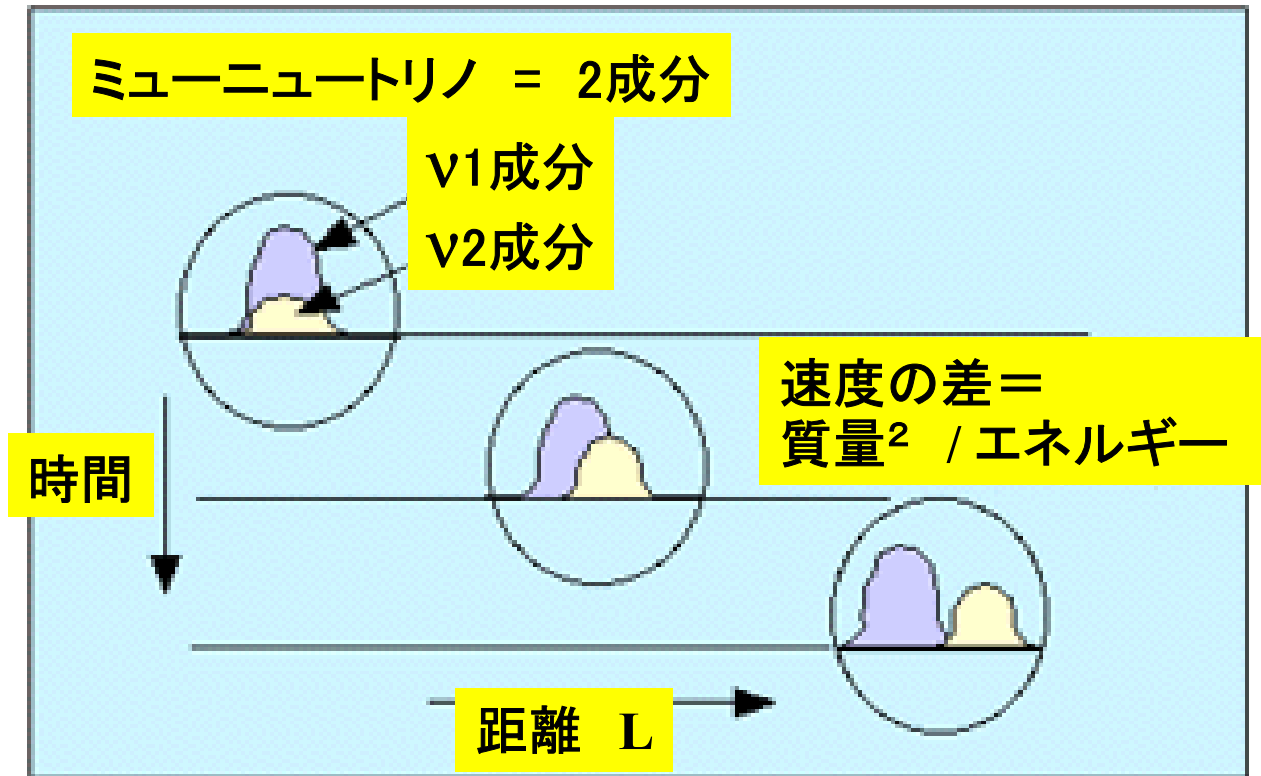
KEK素核研物理第三研究系ニュートリノG

Based on the talk by K.Nishikawa, Neutrino 2002 @ Munich

May 26, 21:50~(JST)

1. Introduction to Introduction
2. Introduction
3. Spectrum measurement at front detector
4. Far/near Extrapolation
5. Oscillation analysis
6. Conclusions

# Neutrino oscillation



ミューニュートリノ = 2成分

$\nu_1$ 成分

$\nu_2$ 成分

時間

速度の差 =  
質量<sup>2</sup> / エネルギー

距離  $L$

$$\text{時間差} = \frac{L}{\text{速度の差}} = \frac{\Delta m^2 \times L}{E}$$

$$P = \sin^2 2\theta \sin^2 \left( 1.27 \frac{\Delta m^2 (\text{eV}^2) L (\text{km})}{E_\nu (\text{GeV})} \right)$$

ミューニュートリノ

$\mu : \tau = 1 : 0$

$|\mu\rangle = -s|1\rangle + c|2\rangle$

$(|\tau\rangle = c|1\rangle + s|2\rangle)$

長距離



•  $|1\rangle = s|\mu\rangle - c|\tau\rangle$

$\mu : \tau = s^2 : c^2$

•  $|2\rangle = c|\mu\rangle + s|\tau\rangle$

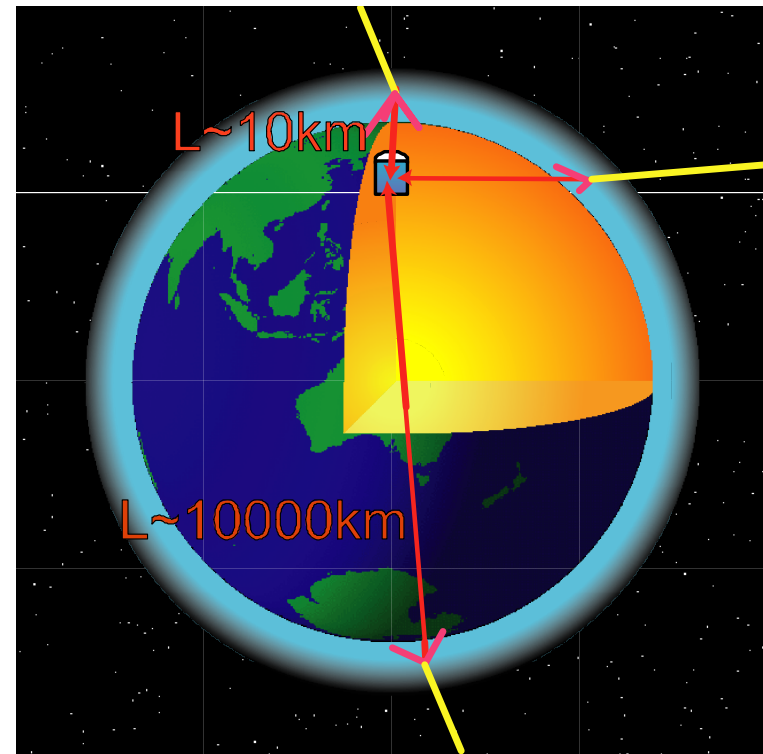
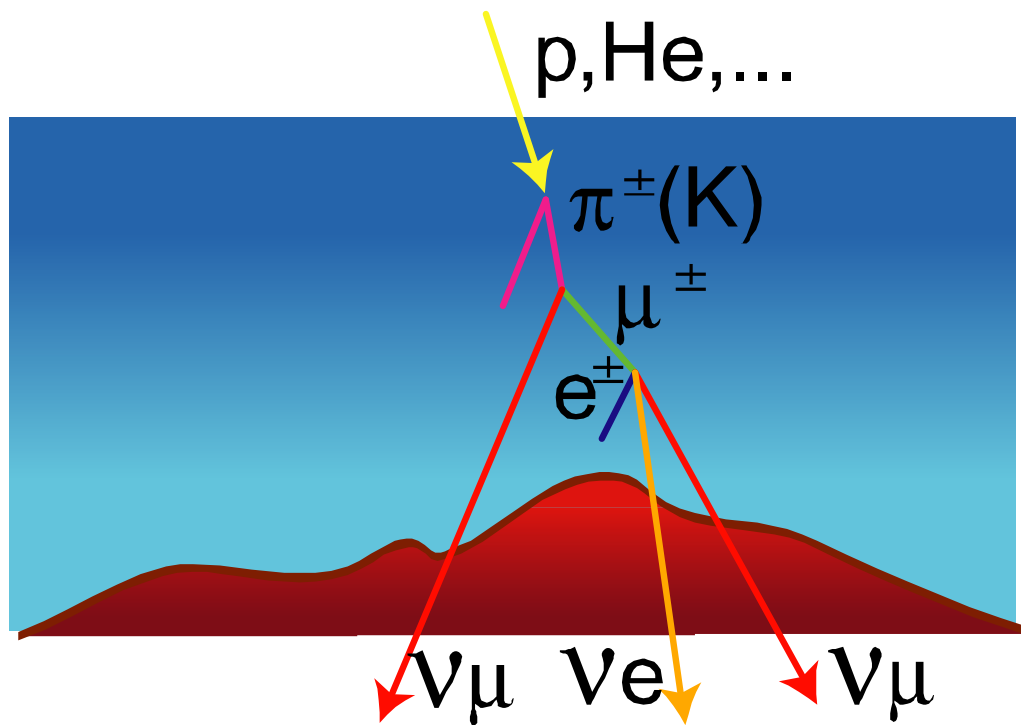
$\mu : \tau = c^2 : s^2$

$\mu : \tau = s^4 + c^4 : 2 * s^2 c^2$

有限な質量<sub>2</sub>

# スーパーカミオカンデによる 大気ニュートリノの観測(1996～)

## 大気ニュートリノ



# Zenith angle distributions (FC+PC+up- $\mu$ )

Neutrino2002  
M. Shiozawa

$\nu_\mu \leftrightarrow \nu_\tau$

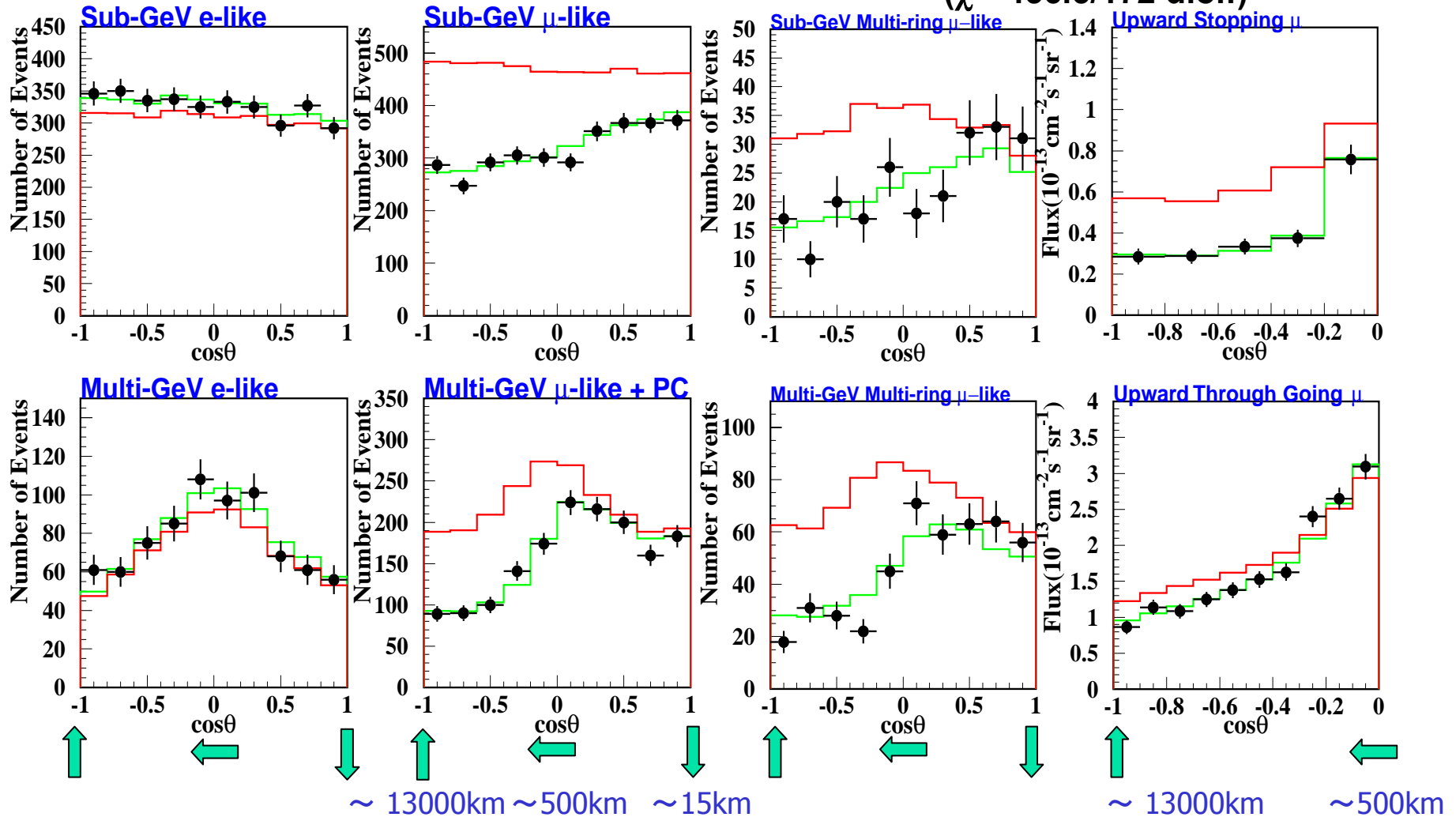
**2-flavor oscillations**

Best fit ( $\Delta m^2 = 2.5 \times 10^{-3} \text{eV}^2$ ,  $\sin^2 2\theta = 1.0$ )

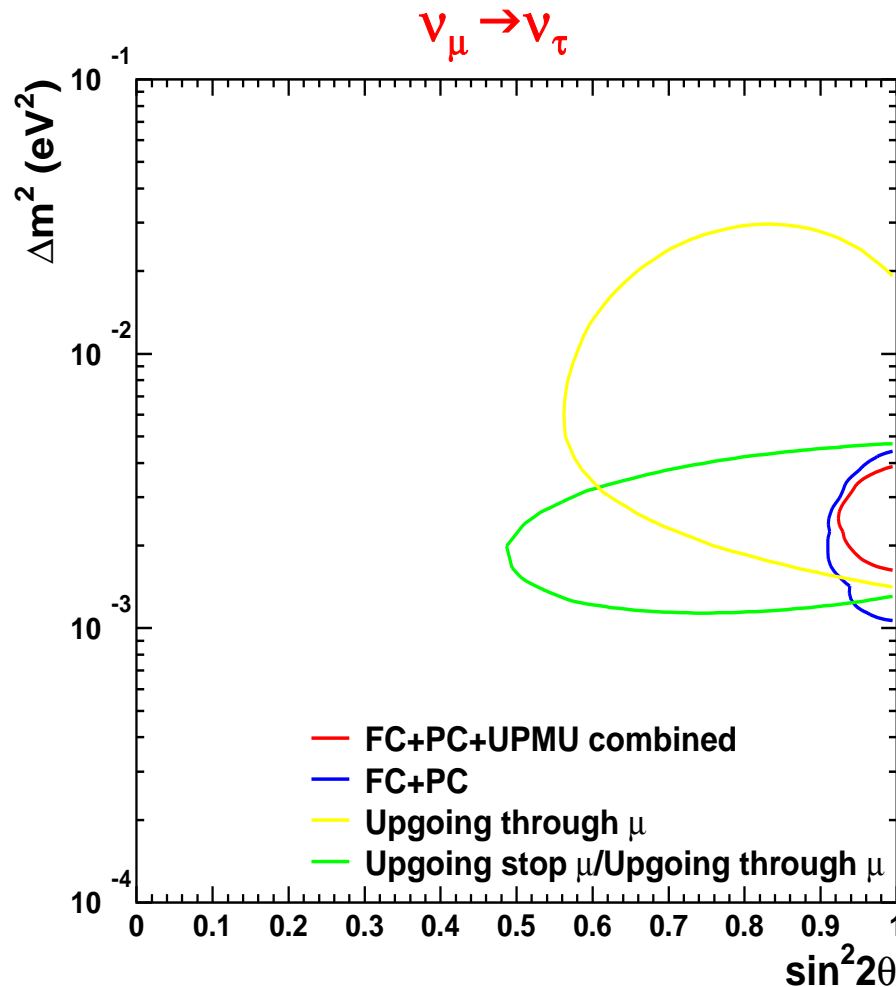
$\chi^2_{\min} = 163.2/170 \text{ d.o.f}$

Null oscillation

( $\chi^2 = 456.5/172 \text{ d.o.f}$ )



# 大気ニュートリノの結果 (Neutrino2002, May26,2002, M.Shiozawa)



## $\nu_\mu \leftrightarrow \nu_\tau$ oscillations

**Best fit**( $\Delta m^2=2.5 \times 10^{-3}$ ,  $\sin^2 2\theta=1.0$ )

$\chi^2_{\min}=163.2/170$  d.o.f)

**No oscillation**

( $\chi^2 = 456.5/172$  d.o.f)

$\Delta m^2 = (1.6 \sim 3.9) \times 10^{-3} \text{eV}^2$

$\sin^2 2\theta > 0.92$  @ 90%CL

# K2K実験(1999年開始)

人工ニュートリノを使って大気ニュートリノの結果を検証

- ①KEKの12GeV陽子シンクロトロンを用いて $\nu_{\mu}$ ビーム( $\sim 1\text{GeV}$ )を生成。
  - ②250km離れたスーパーカミオカンデ(SK)に向けて撃ち出す。
  - ③SKでKEKからのニュートリノを検出し、その数、種類、エネルギー分布を測定する。
- ミューオンニュートリノ消失実験( $\nu_{\mu}$  disappearance)
  - (電子ニュートリノ出現実験:  $\nu_e$  appearance)

# ミューオンニュートリノ消失実験 ( $\nu_\mu$ disappearance)

生成

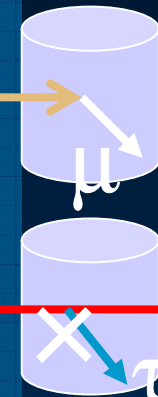
$\sim 1\text{GeV } \nu_\mu$

生成

$1\text{GeV } \nu_\tau$



$$p(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta \cdot \sin^2 \left( 1.27 \frac{\Delta m^2 [\text{eV}^2] \cdot L [\text{km}]}{E_\nu [\text{GeV}]} \right)$$



- ニュートリノ反応、識別



- 生成されるレプトンで識別(eなら $\nu_e$ ,  $\mu$ なら $\nu_\mu$ ,...)

- 反応閾値

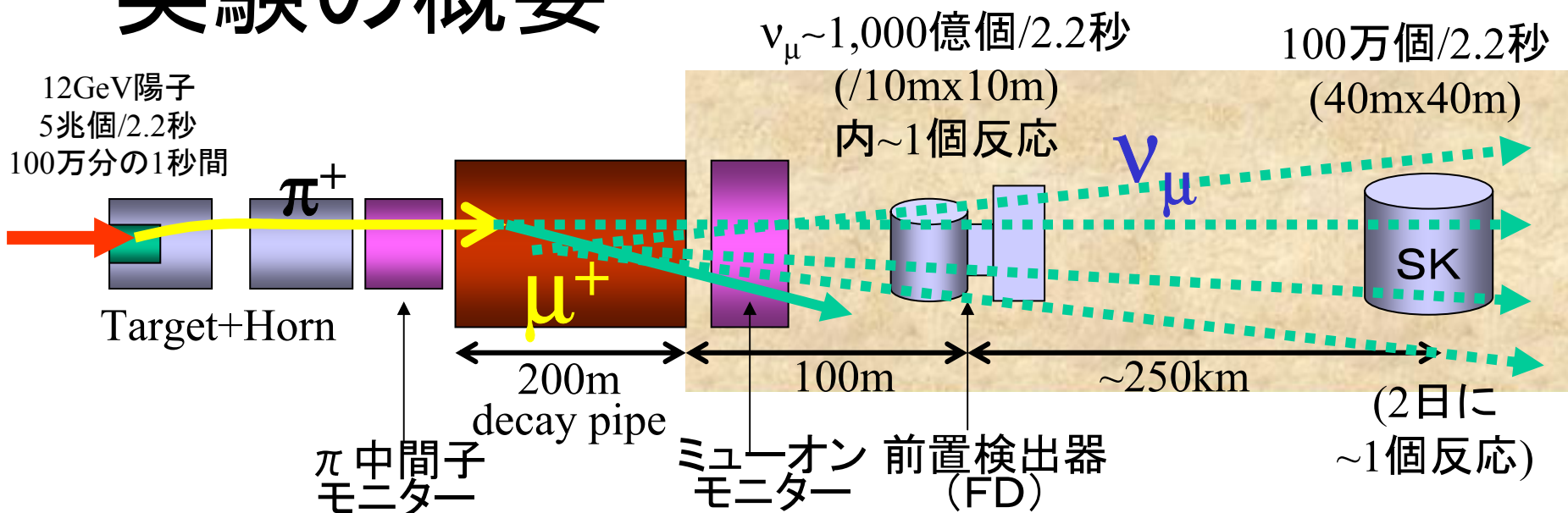
$\nu_\mu \rightarrow \mu: 110\text{MeV}$ ,  $\nu_\tau \rightarrow \tau: 3.5\text{GeV}$  (生成レプトンの質量による)

- $1\text{GeV } \nu_\mu$ が振動して $\nu_\tau$  ( $1\text{GeV}$ ): 反応できない

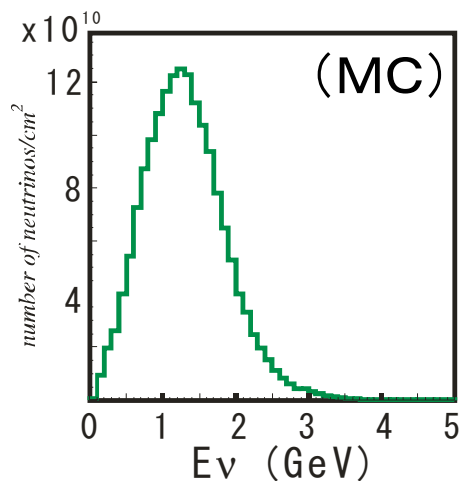
- 信号

- 振動していない場合に比べ、ニュートリノ反応の数が減る。
- 残った $\nu_\mu$ のエネルギーはニュートリノ振動に特有の分布を示す

# 実験の概要

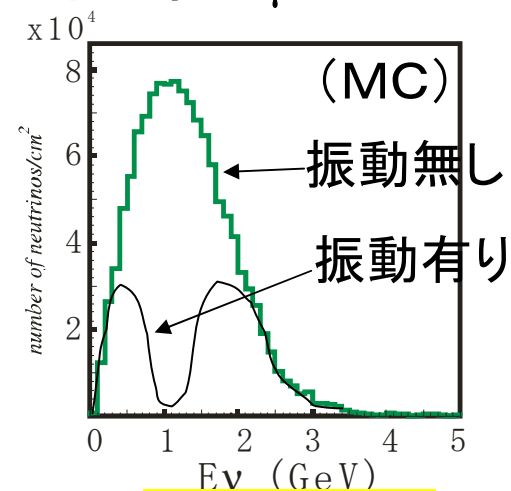


生成時の $\nu_\mu$ スペクトル

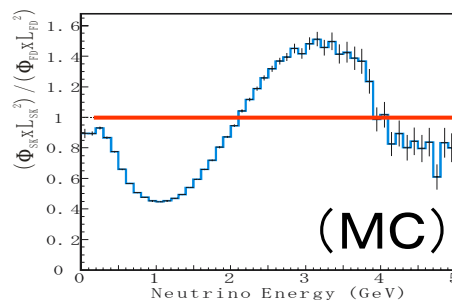


前置検出器で測定

SK到達時の $\nu_\mu$ スペクトル



観測と比較

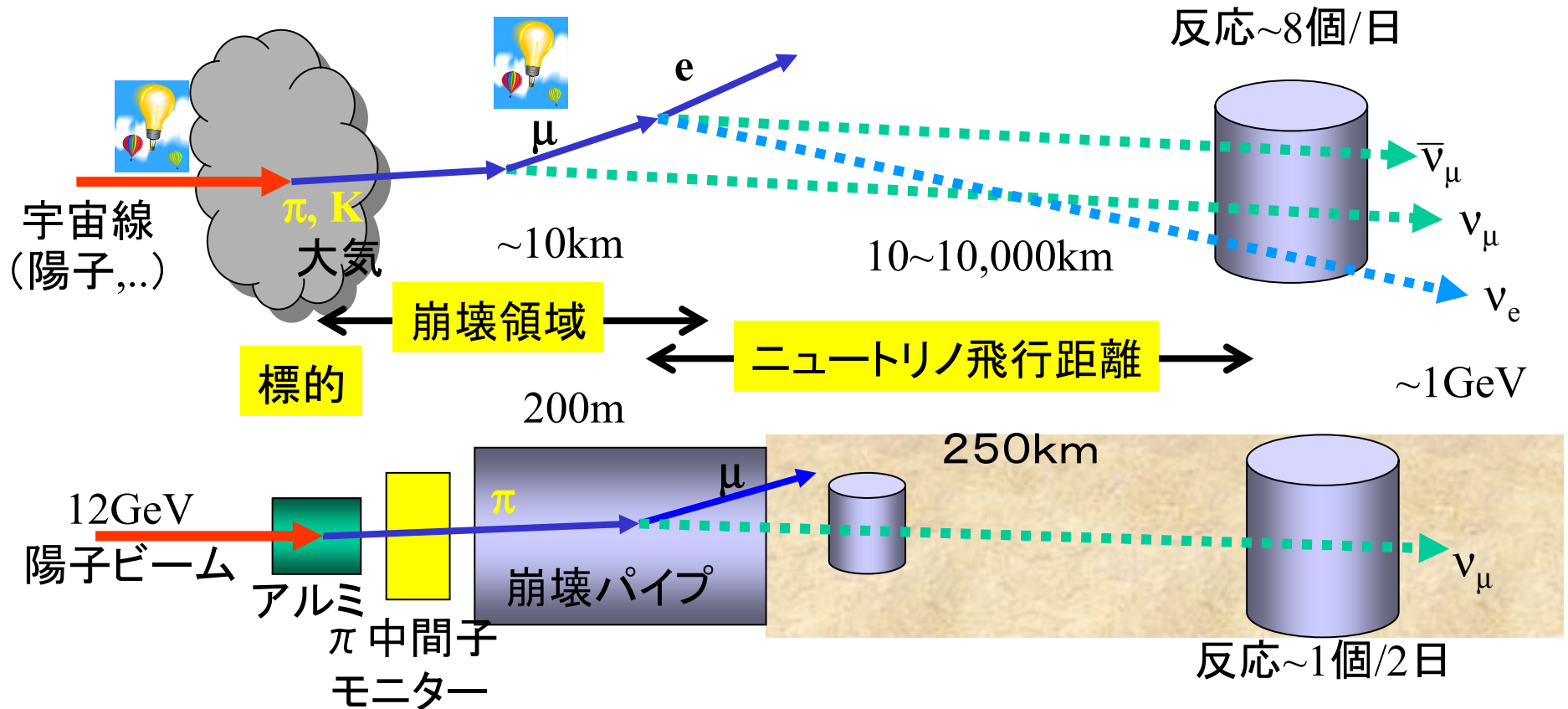


スペクトルの比  $\neq 1$

$\pi$ 中間子モニターで測定



# 大気ニュートリノ vs K2K



- K2K: ~100%  $\nu_\mu$ , Atm:  $\nu_\mu:\nu_e \sim 2:1$
- K2K: ニュートリノの飛行距離決まっている。(250km) → 特に $\Delta m^2$ 測定に高感度
- K2K: ニュートリノの方向決まっている。  
→一部の反応生成粒子から親ニュートリノのエネルギー算出可能
- K2K: 生成直後の $\pi$ 中間子、振動前ニュートリノの(数、エネルギー分布)直接測定
- Atm: 反応数大
- Atm: 広い $\Delta m^2$ 探索可能“discovery machine”, 加速器実験: 狙いを絞った“精密実験”

# **K2K results**

## **Results of Total Event Rate & Spectrum Shape Analysis**

**K2K collaboration**

**K.Nishikawa**

**Kyoto University**

**NEUTRINO 2002, May 26,2002**

**Munich, Germany**

# **K2K Collaboration**

**High Energy Accelerator Research Organization(KEK)  
Institute for Cosmic Ray Research(ICRR), University of Tokyo**

**Kobe University**

**Kyoto University**

**Niigata University**

**Okayama University**

**Tokyo University of Science**

**Tohoku University**

**Chonnam National University**

**Dongshin University**

**Korea University**

**Seoul National University**

**Boston University**

**University of California, Irvine**

**University of Hawaii, Manoa**

**Massachusetts Institute of Technology**

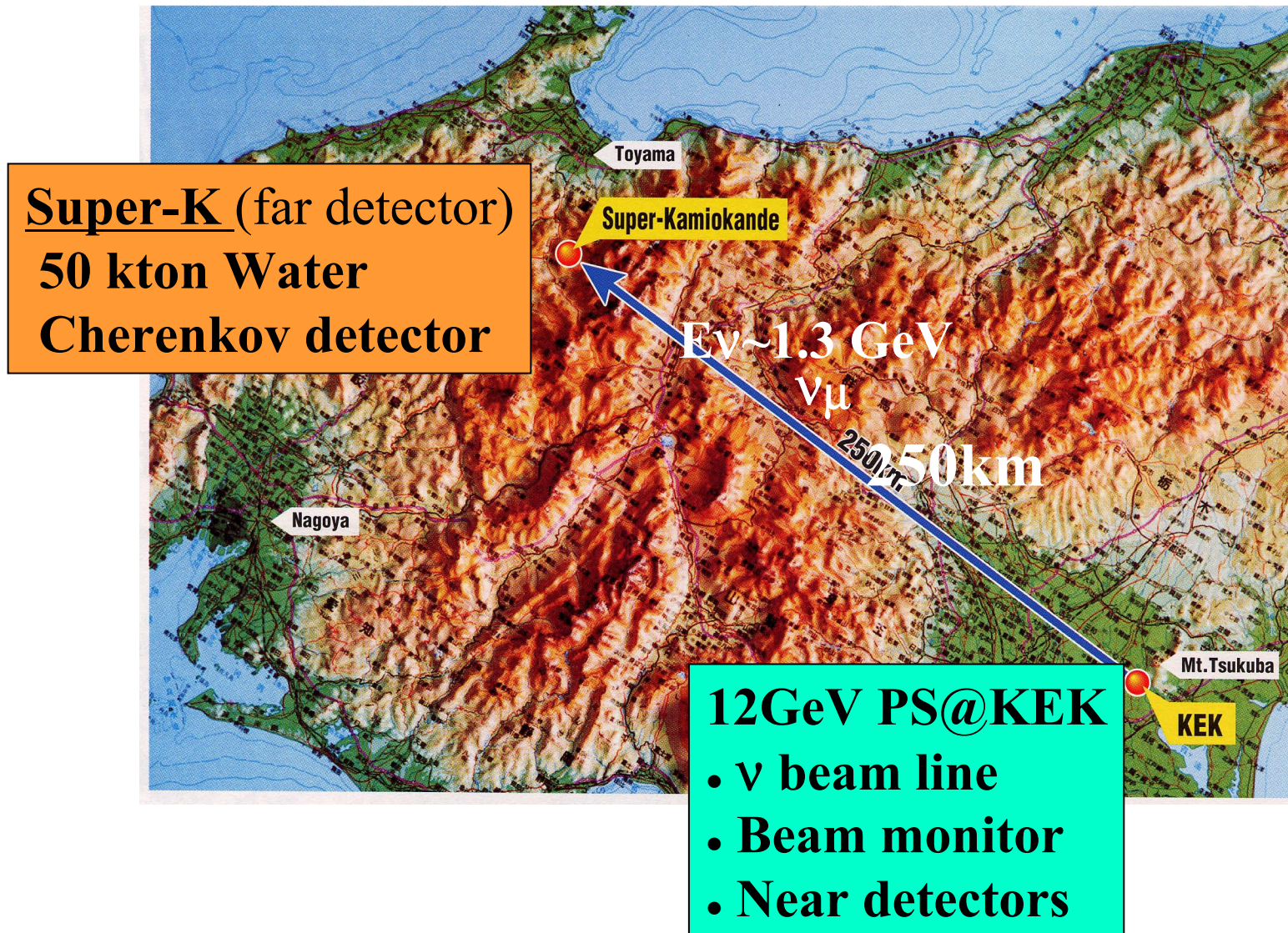
**State University of New York at Stony Brook**

**University of Washington at Seattle**

**Warsaw University**

**Solton Institute**

# KEK to Kamioka Neutrino Oscillation Experiment



# Principle of K2K

**Fixed distance**

**( $E_\nu \sim 1.3$  GeV,  $L=250$ km)**

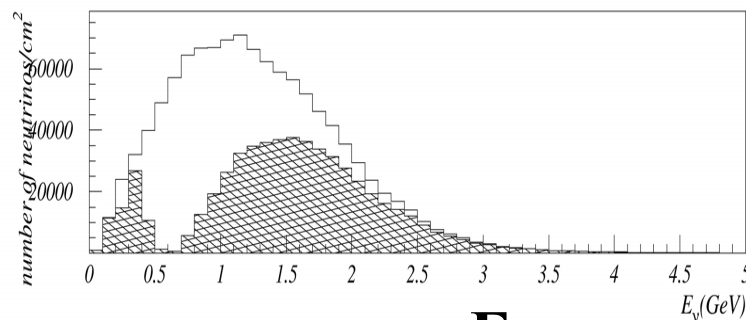
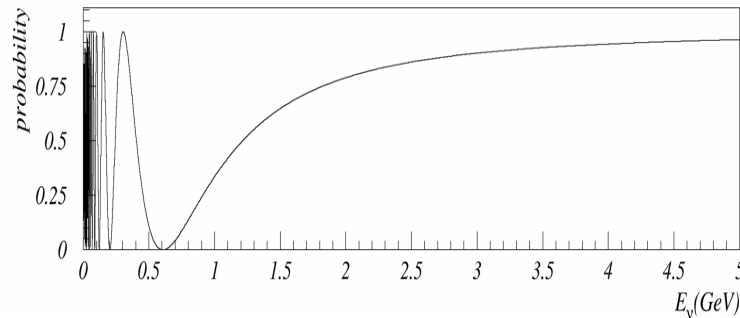
**(99%  $\nu_\mu$ ,  $\sigma_\tau \ll \sigma_\mu$ )**

$$\text{prob.} = \sin^2 2\theta \cdot \sin^2 \left( \frac{1.27 \Delta m^2 L}{E_\nu} \right)$$

**Results**

- **Reduction of events**
- **Spectrum distortion**

**Neutrino Oscillation ( $\Delta m^2=0.003\text{eV}^2$ )**



**$E_\nu$**

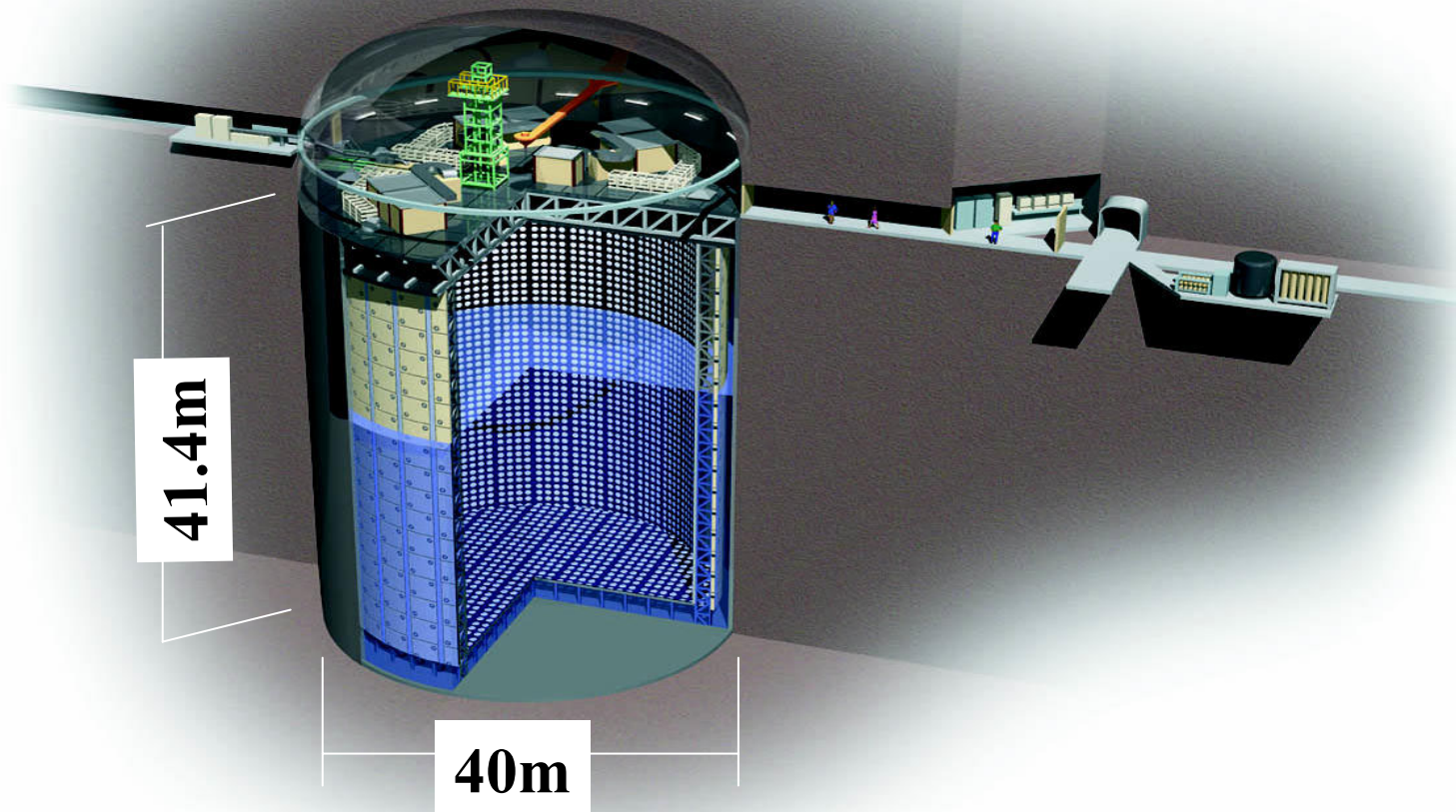
- **Neutrino beam at near site**
- **Near to Far extrapolation**
- **Rate and spectrum shape @ SK**

(April 1996 commissioned)

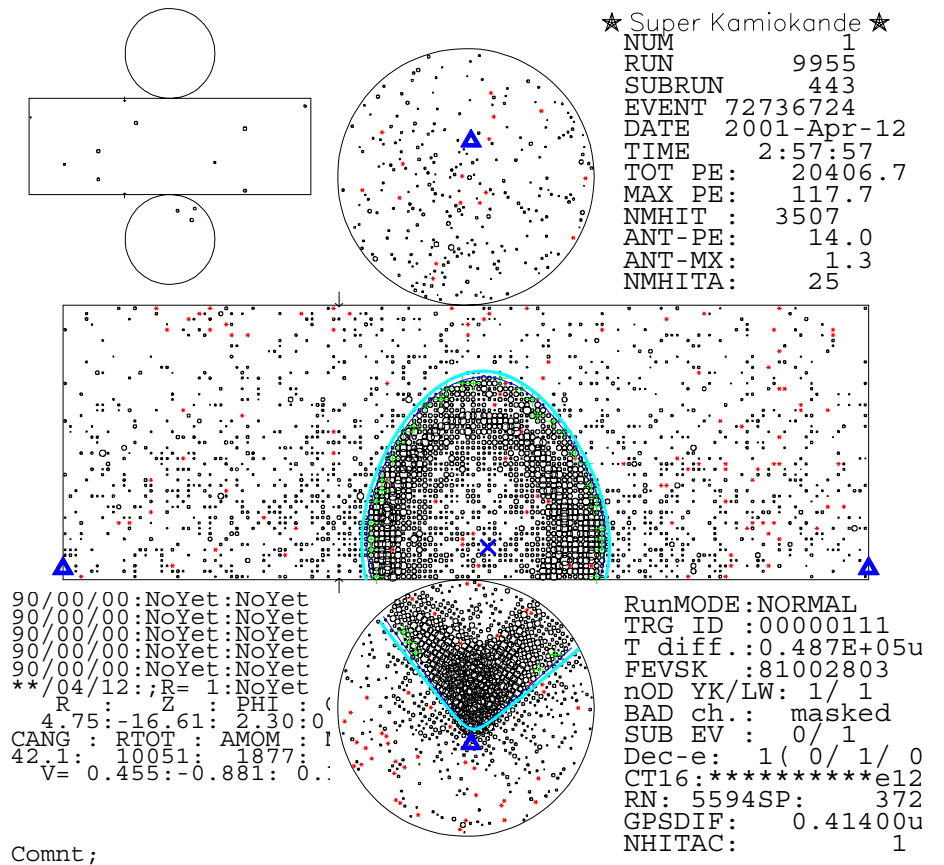
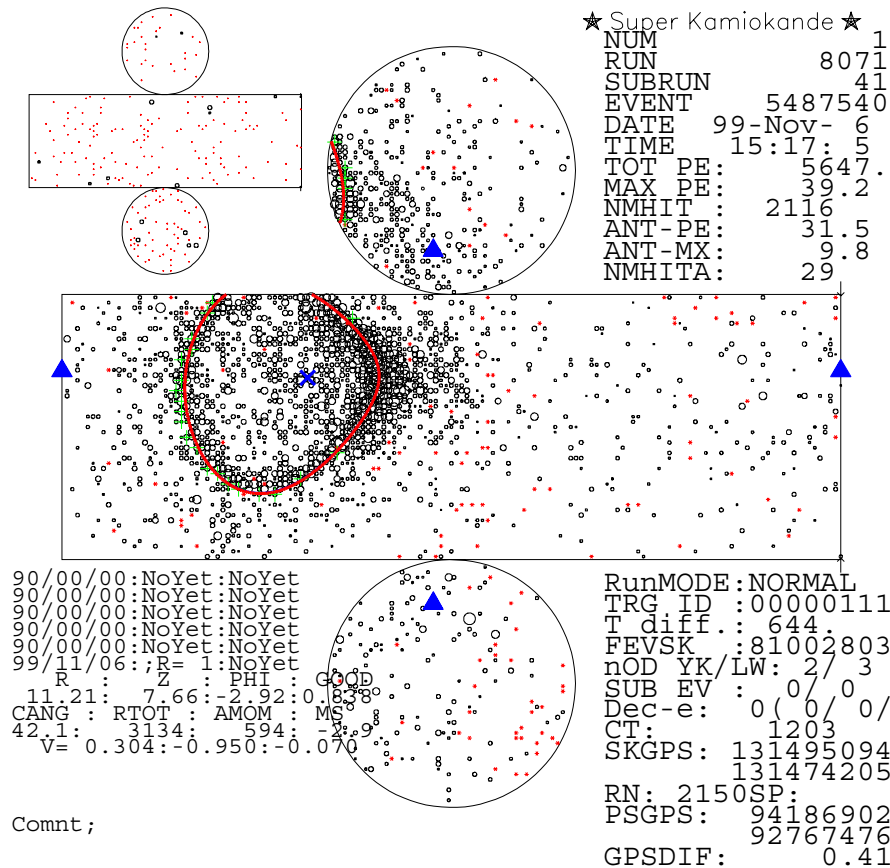
# Super-Kamiokande

50,000 ton water Cherenkov detector (22.5 kton fiducial volume)

Optically separated **INNER** and **OUTER** detector



# e-like and $\mu$ -like events in Super-Kamiokande



**Total rate with low threshold (>30MeV)**  
**Identification of  $\mu$  (1R $\mu$ ) , e (1Re)**

$\mu$ -monitor  
Front (Near) Detector

direction ( $\pi \rightarrow \mu$ )  
direction ( $\nu$ )  
spectrum, rate

12 GeV PS  
>  $5 \times 10^{13}$  ppp  
2.2sec/pulse

North  
Counter  
Hall

Target/Double Horn  
 $\sim 20 \times$  flux

Front detector

$\mu$ -monitor



12 GeV PS

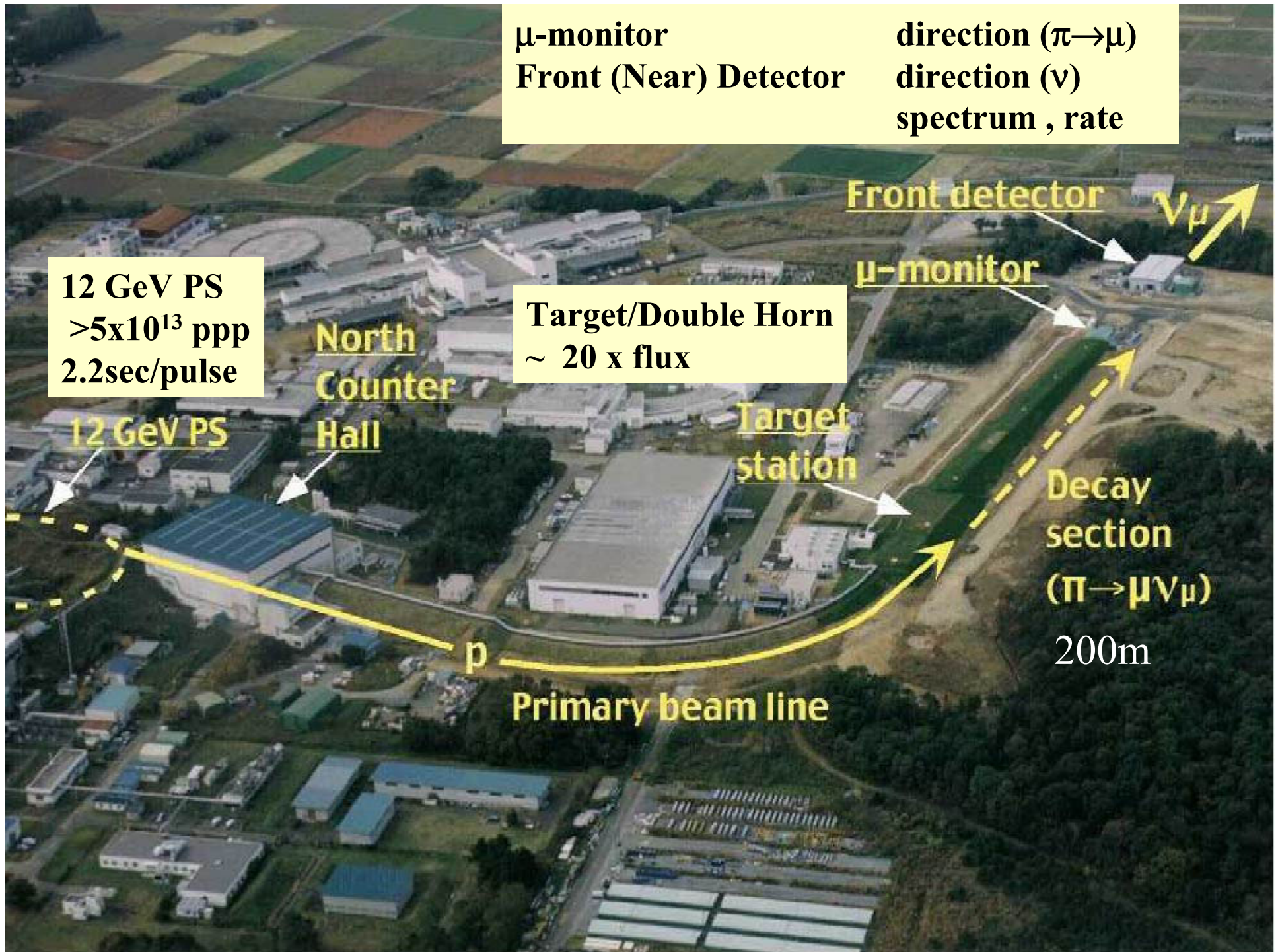
Target  
station

Decay  
section  
( $\pi \rightarrow \mu \nu_\mu$ )

200m

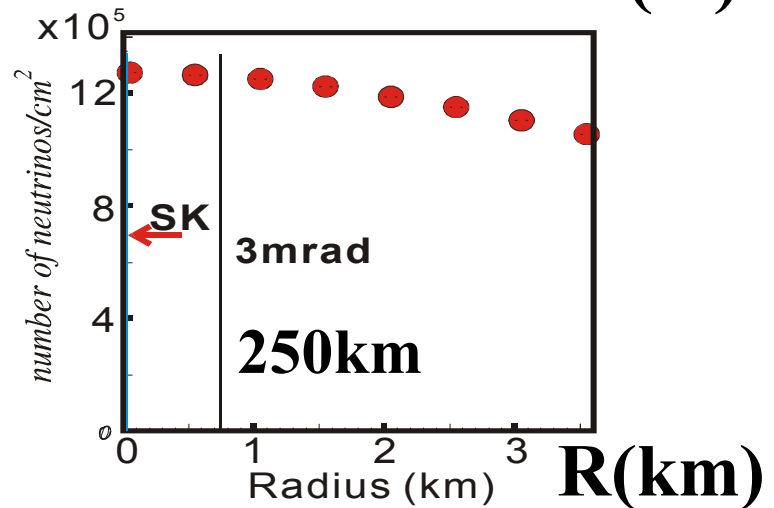
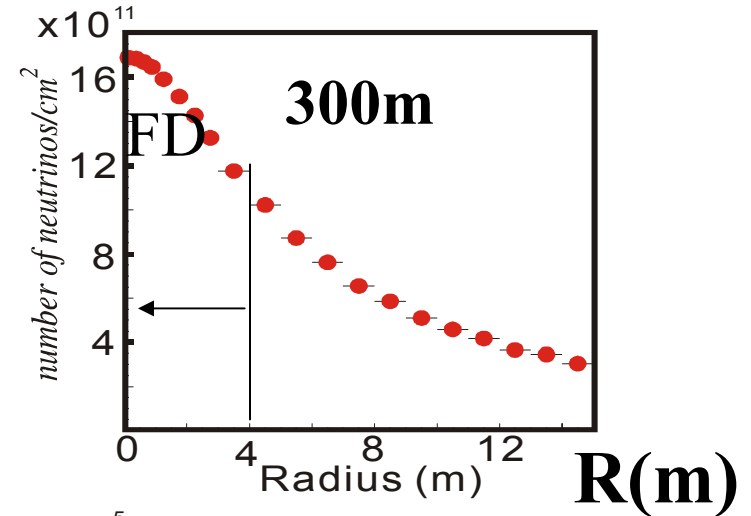
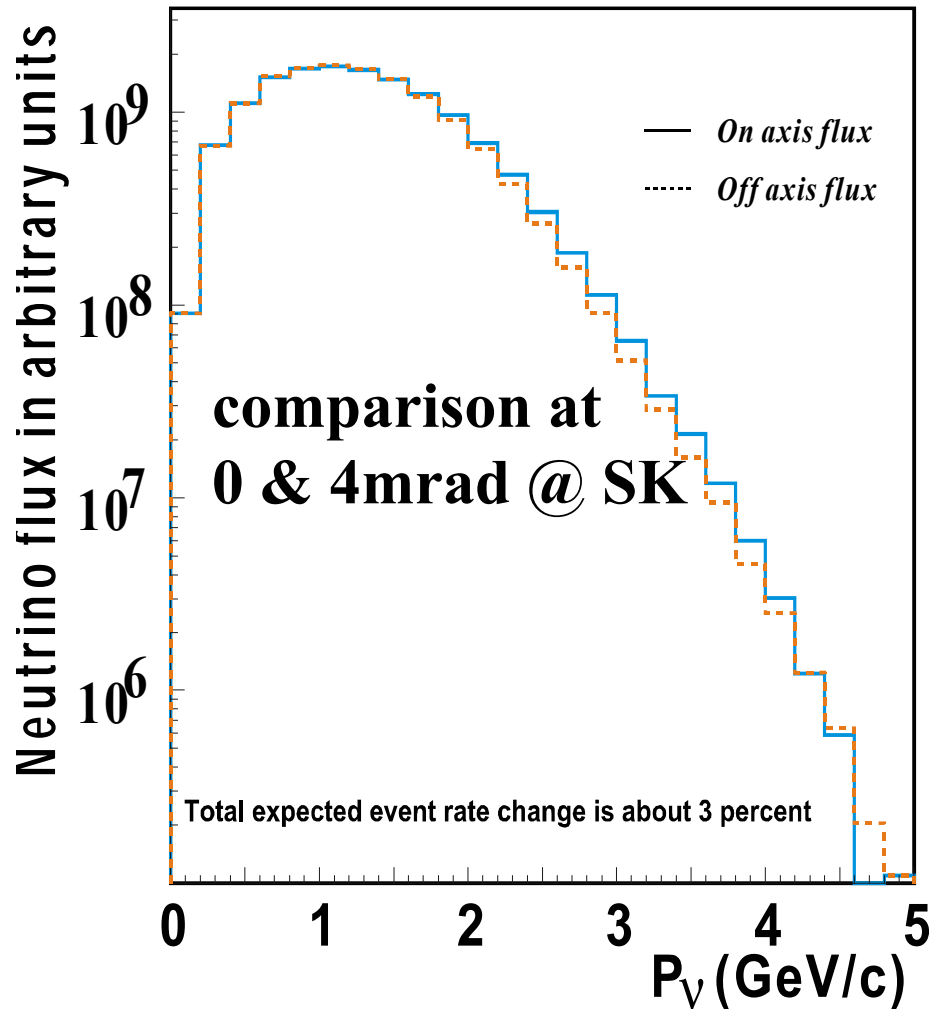
p

Primary beam line





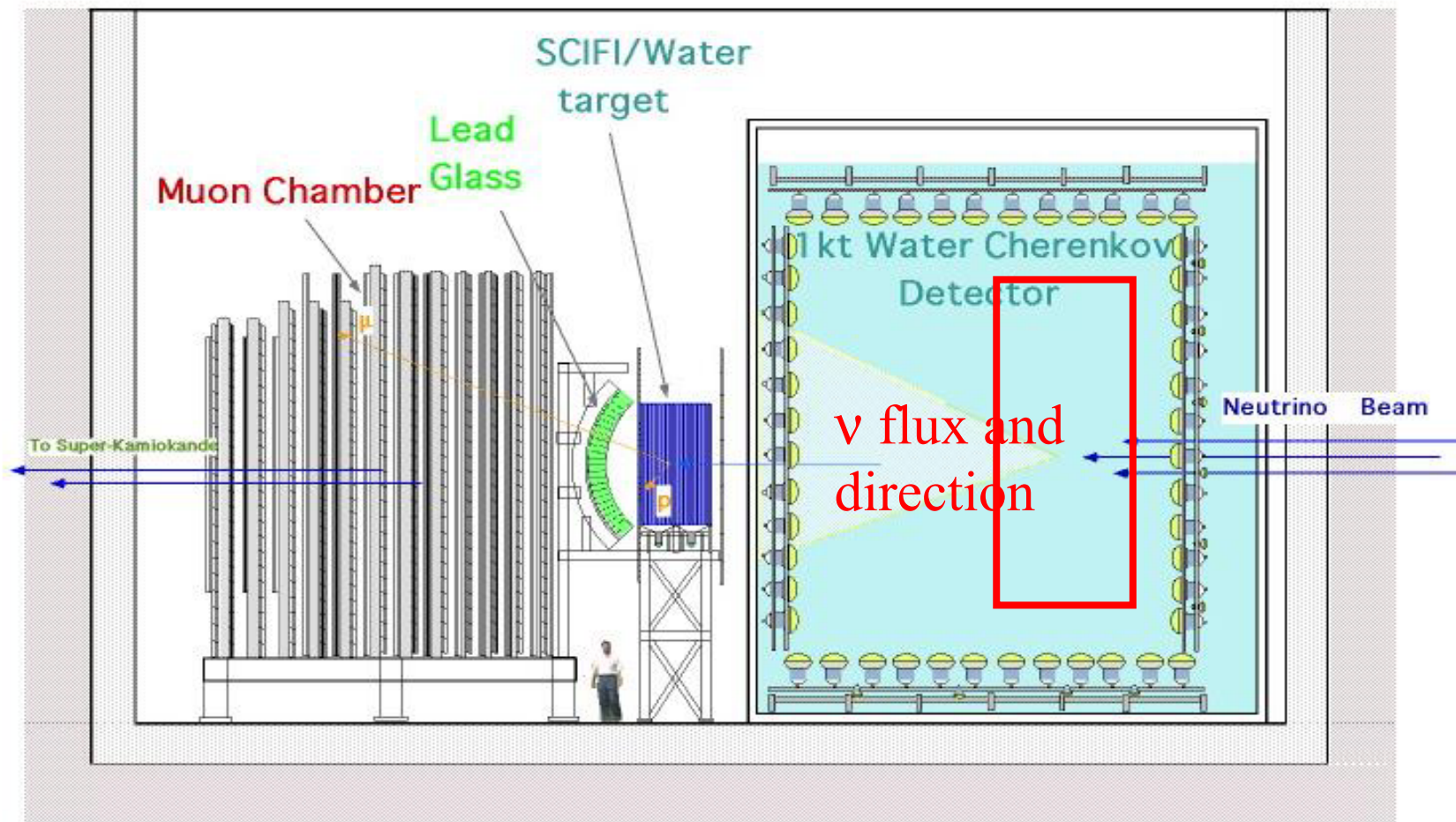
# Expected (MC) Neutrino Spectra and Radial Distributions at 300m/250km



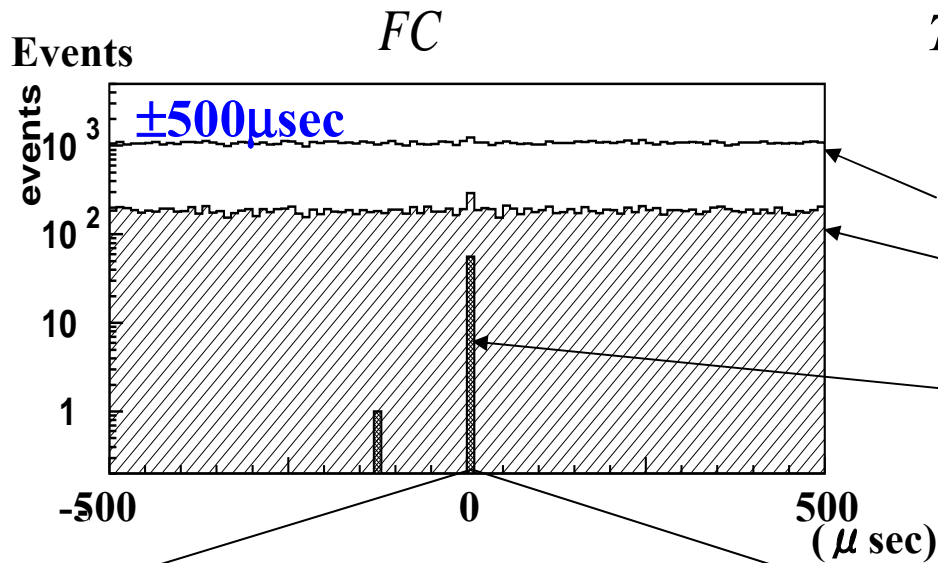
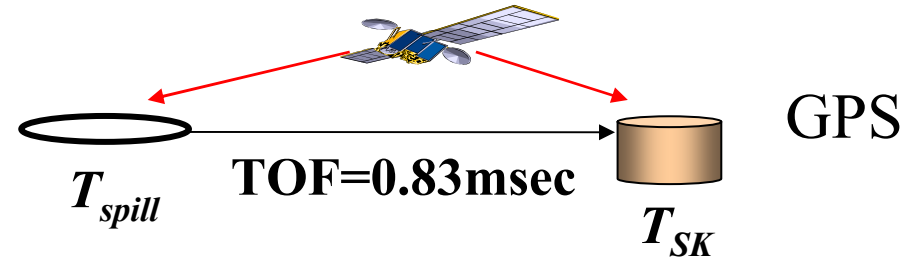
1km(4mr) off axis @ SK no change in rate and spectrum

# Near Detectors

- At KEK: **1kt Water Cherenkov detector (KT) 25 ton fiducial**  
+ **Water tube sandwiched Scintillation fiber detector (SciFi) 6 ton**  
+ **Muon range detector(MRD) 700 ton direction by  $\nu$  & stability**  
+ **Lead glass detector (LG)**



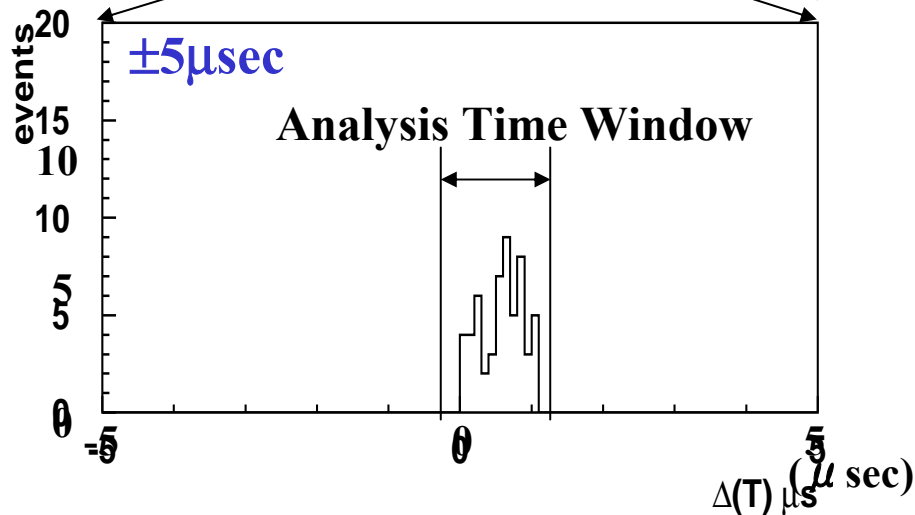
# Super-Kamiokande Event Selection



Not a Decay Electron Event

More than 20MeV Deposited Energy

**No Activity in Outer Detector**  
**Event Vertex is in Fiducial Volume**  
**More than 30MeV Deposited Energy**



**56 Events**

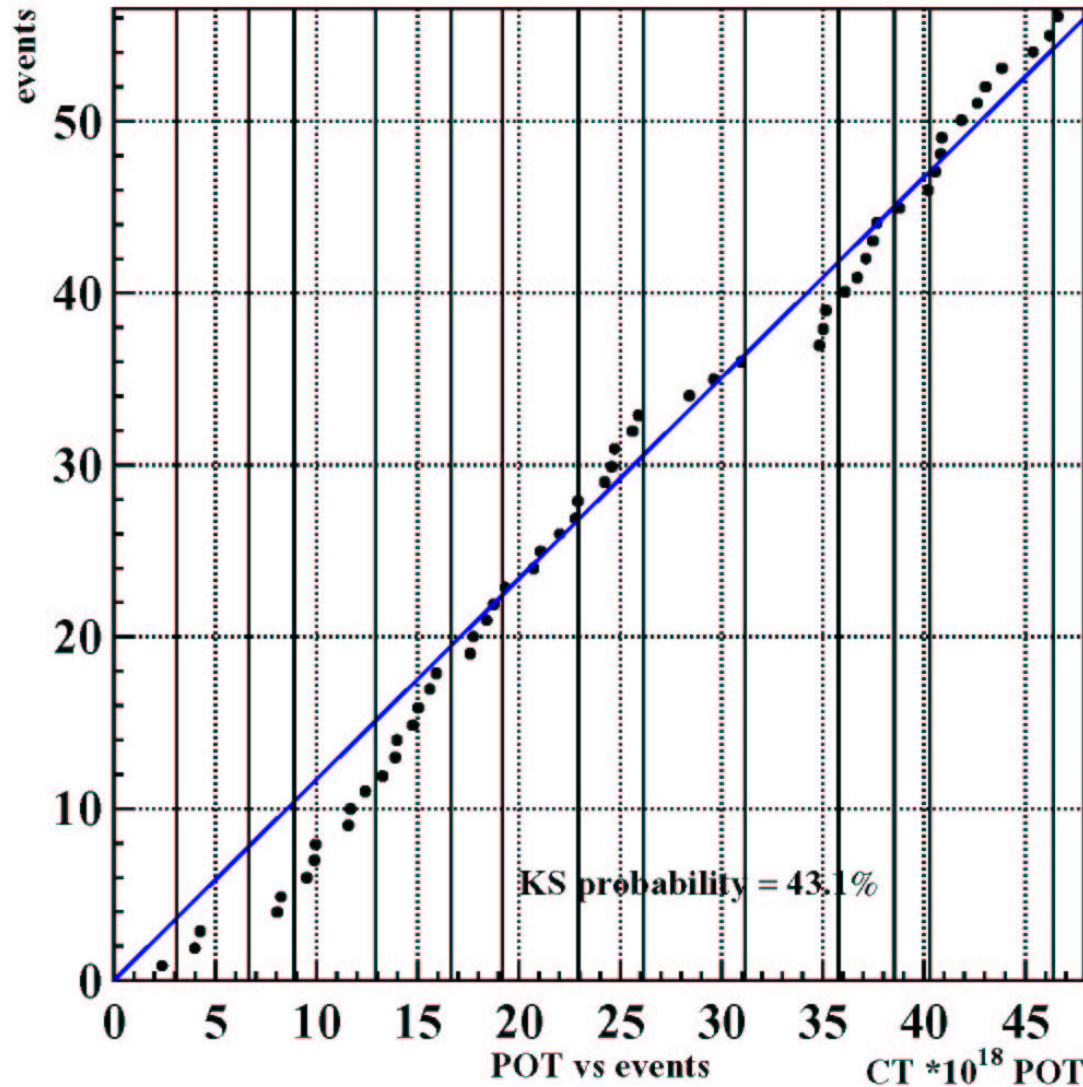
in

$-0.2 < T_{SK} - T_{spill} - TOF < 1.3\mu\text{sec}$

Analysis Time Window

(B.G.  $\sim 10^{-3}$  in  $1.5\mu\text{s}$  Window)

# Event POT distribution



KS test probability

**43%**

## Summary of K2K results in last year

- Accumulated  $4.8 \times 10^{19}$  POT @ SK from Jun '99 to July '01.
- Neutrino beam is well under control
  - direction  $< 1$  mrad. (pulse by pulse  $\pi \rightarrow \mu$ ,  $\nu$  interaction vertex profile)
  - Stability of  $E_\mu$  spectrum from  $\nu$  interaction
- # of fully contained events in fiducial volume (FCFV) @ SK  
Observed: **56**, Expected with null oscillation **80**  $^{+7.3}_{-8.0}$   
Probability of null oscillation  $< 3\%$



- Full & improved error estimates
- Spectrum shape analysis

# Flow of Neutrino Oscillation Analysis

Observed  $(p_\mu, \theta_\mu)$  distributions at Near Detectors

↓ *v Int. Model*

Neutrino Spectrum at Near detector  $\phi_{near}(E\nu)$ ,

↓

Far/Near Extrapolation vs  $E\nu$   $R_{FN}(E\nu)$

Neutrino Spectrum w/o oscillation at SK  $\phi_{SK}(E\nu)$

$\phi_{SK}(E\nu) \otimes$  Oscillation  $(\sin^2 2\theta, \Delta m^2) \otimes$  *Int. Model*

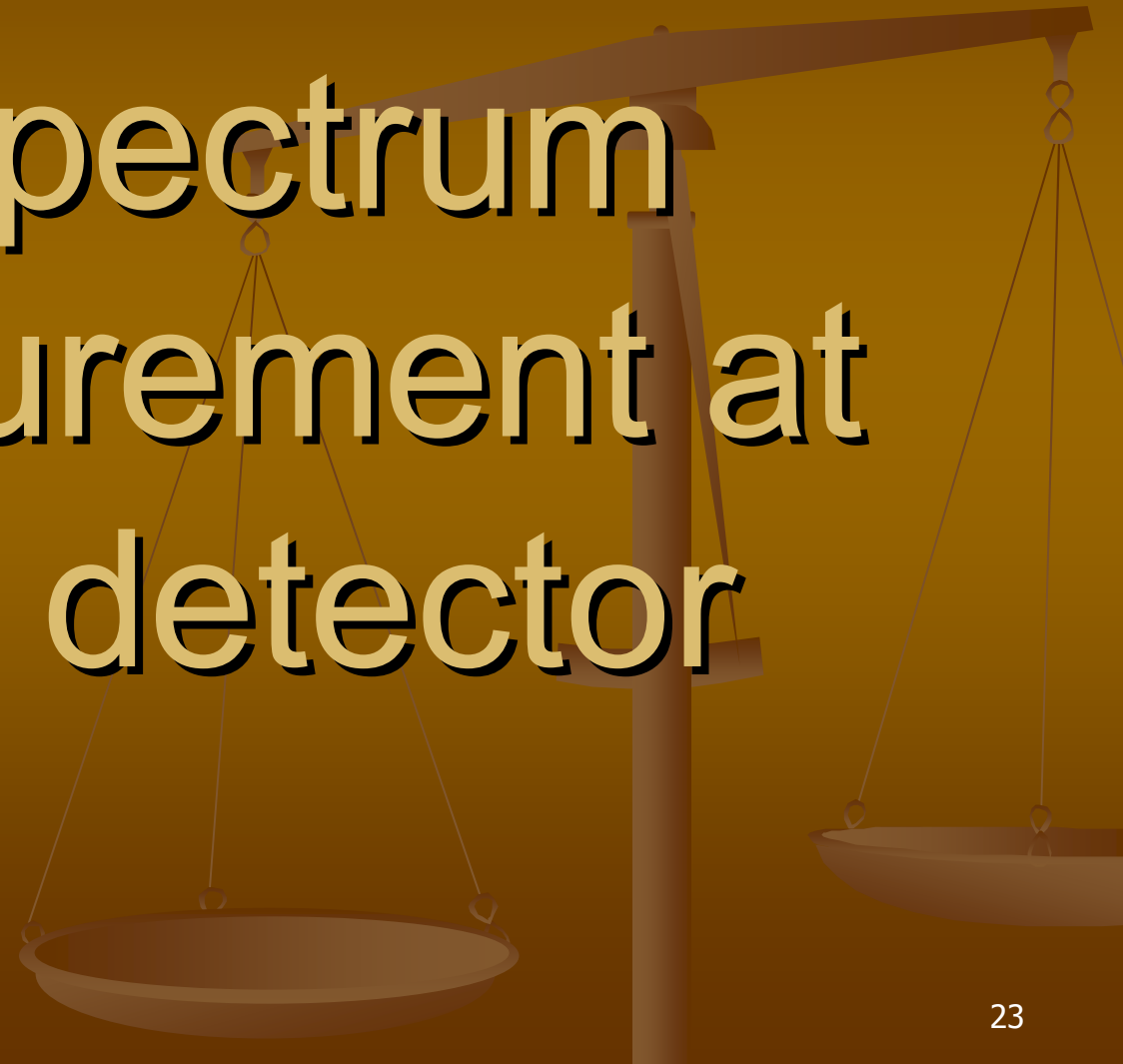
## Prediction

- $N_{SK}(\text{exp't})$  : Expected no. of SK events
- $S_{SK}(E_\nu^{rec})$  :  $1R\mu$   $E_{rec}$  distribution(shape)

## SK observation

- $N_{SK}(\text{obs})$
- $1R\mu$   $E_{rec}$  distribution

Maximum Likelihood Fit in  $(\sin^2 2\theta, \Delta m^2)$



$\nu_{\mu}$  spectrum  
measurement at  
front detector

# 1R $\mu$ events in water Cherenkov detector

$\nu_{\mu} + n \rightarrow \mu + p$

$\mu^{-}$   
( $E_{\mu}, p_{\mu}$ )

$\theta_{\mu}$

$\nu$

$p$

- ✧ CC QE
- ✧ ~100% efficiency for  $N_{SK}$
- ✧ can reconstruct  $E_{\nu}$

$\nu_{\mu} + n \rightarrow \mu + p + \pi$

$\mu^{-}$   
( $E_{\mu}, p_{\mu}$ )

$\theta_{\mu}$

$\nu$

$\pi's$   $p$

- ✧ CC nQE
- ✧ ~100% efficiency for  $N_{SK}$
- ✧ Bkg. for  $E_{\nu}$  measurement

$\nu_{\mu} + n \rightarrow \nu + p + \pi's$

$\nu$

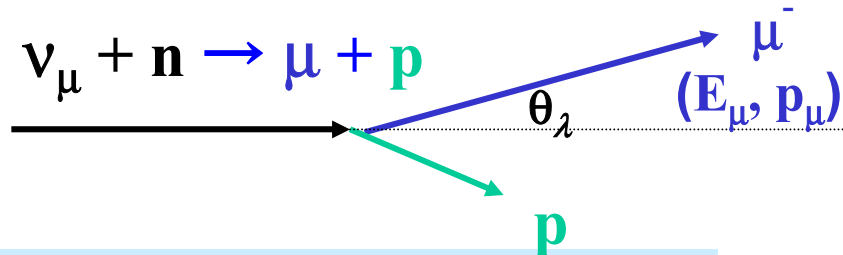
$\pi's$   $p$

- ✧ NC
- ✧ ~40% efficiency for  $N_{SK}$



# Neutrino Energy $E_\nu$ Reconstruction

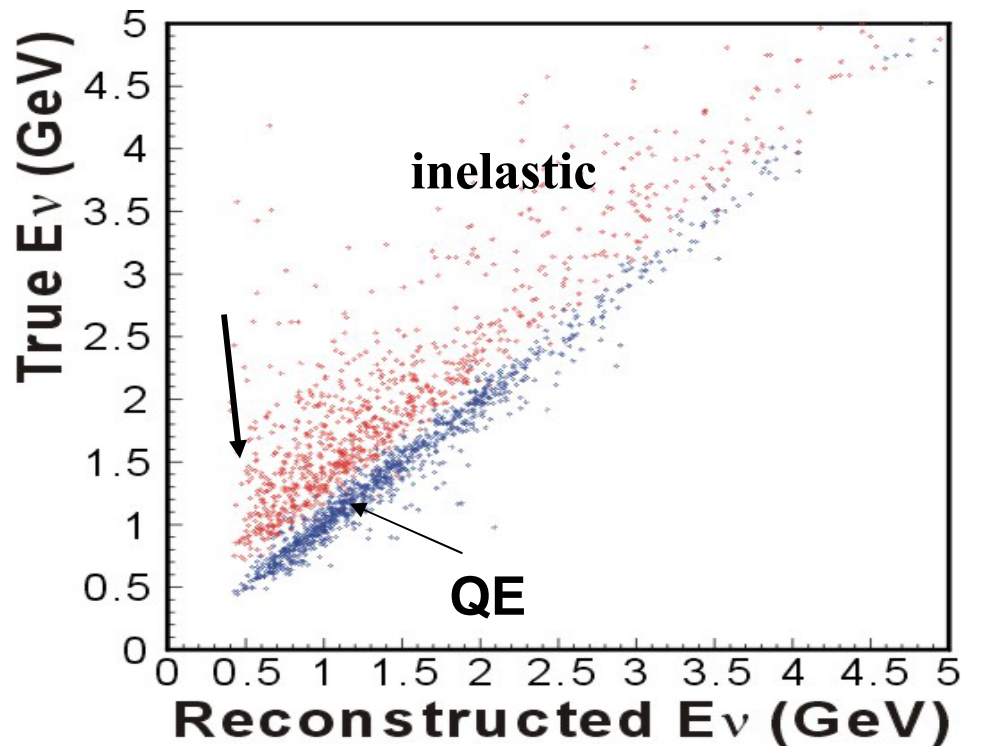
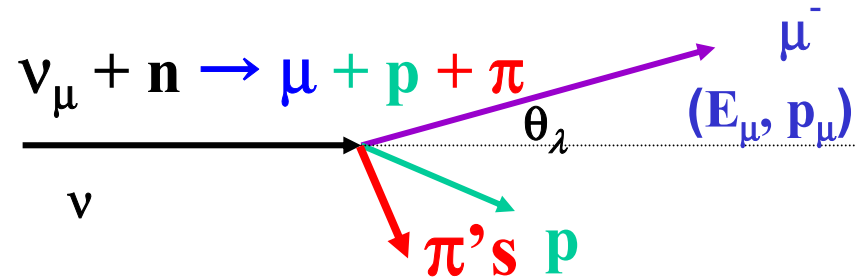
CC quasi elastic (QE)



$$E_\nu = \frac{m_N E_\mu - m_\mu^2 / 2}{m_N - E_\mu + p_\mu \cos \theta_\mu}$$

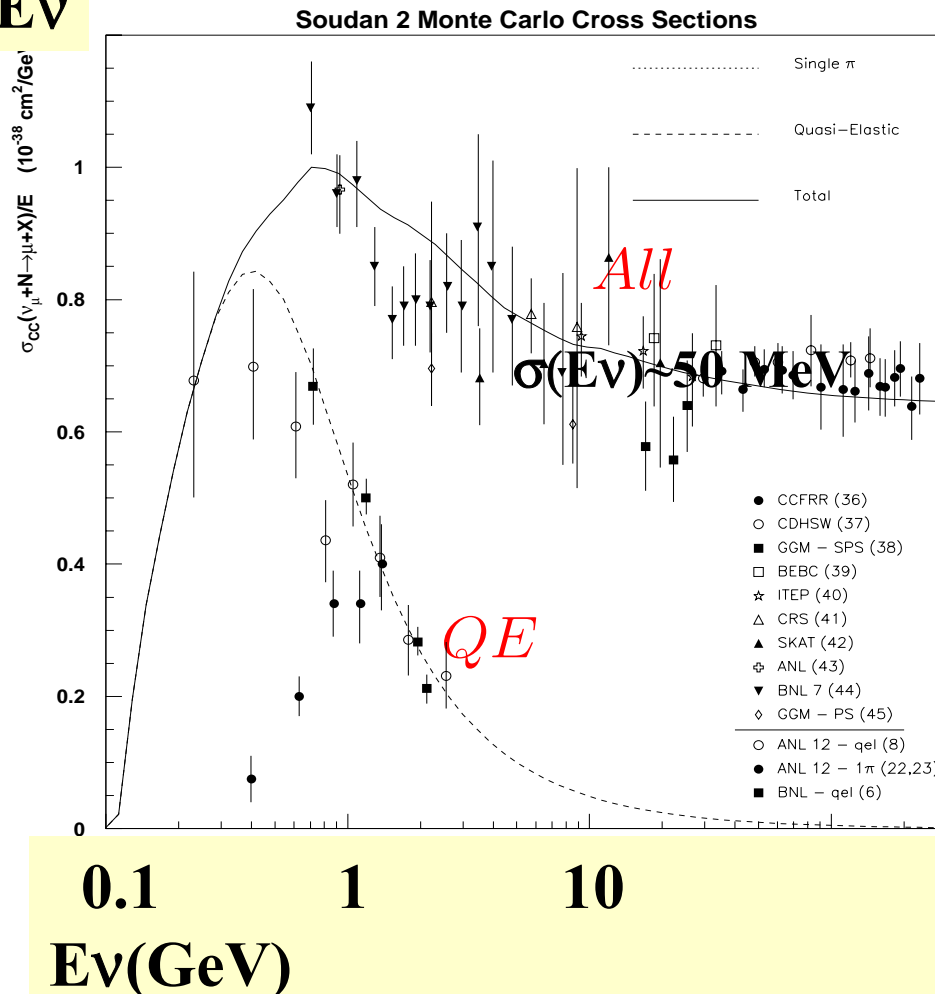
Rate( $E_\nu$ , Near)  $\rightarrow$   $\phi(E_\nu$ , Near)  
 $\uparrow$   
 $\sigma(\text{QE}), \sigma(\text{nonQE})$

CC inelastic



# CC Quasi Elastic(QE) and Other Processes(nQE)

$\sigma/E\nu$



Not well known

Used Parameters

$MA(QE)=1.11\text{GeV}$

$MA(1\pi)=1.21 \text{ GeV}$

Coherent  $\pi$  : Marteau et.al.

Multi- $\pi$ : use hep-ex/0203009

Checked

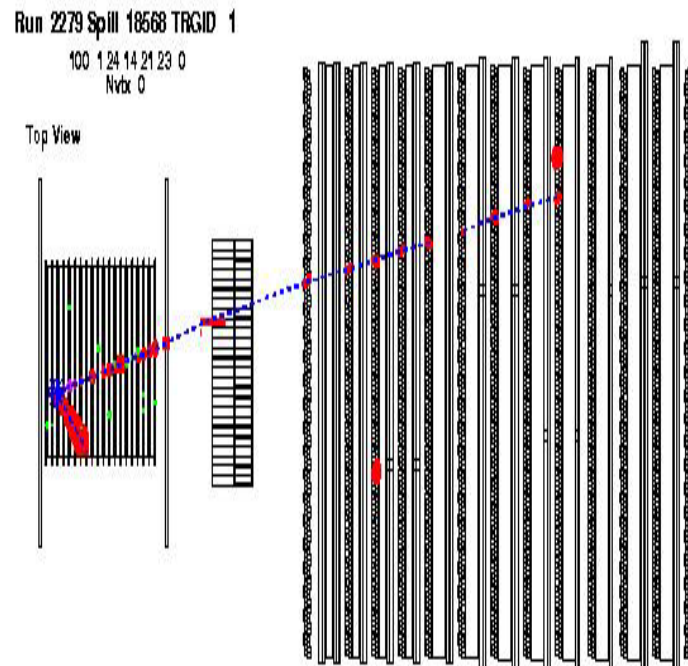
$MA(QE)=1.01-1.11$

$MA(1p)=1.01-1.51$

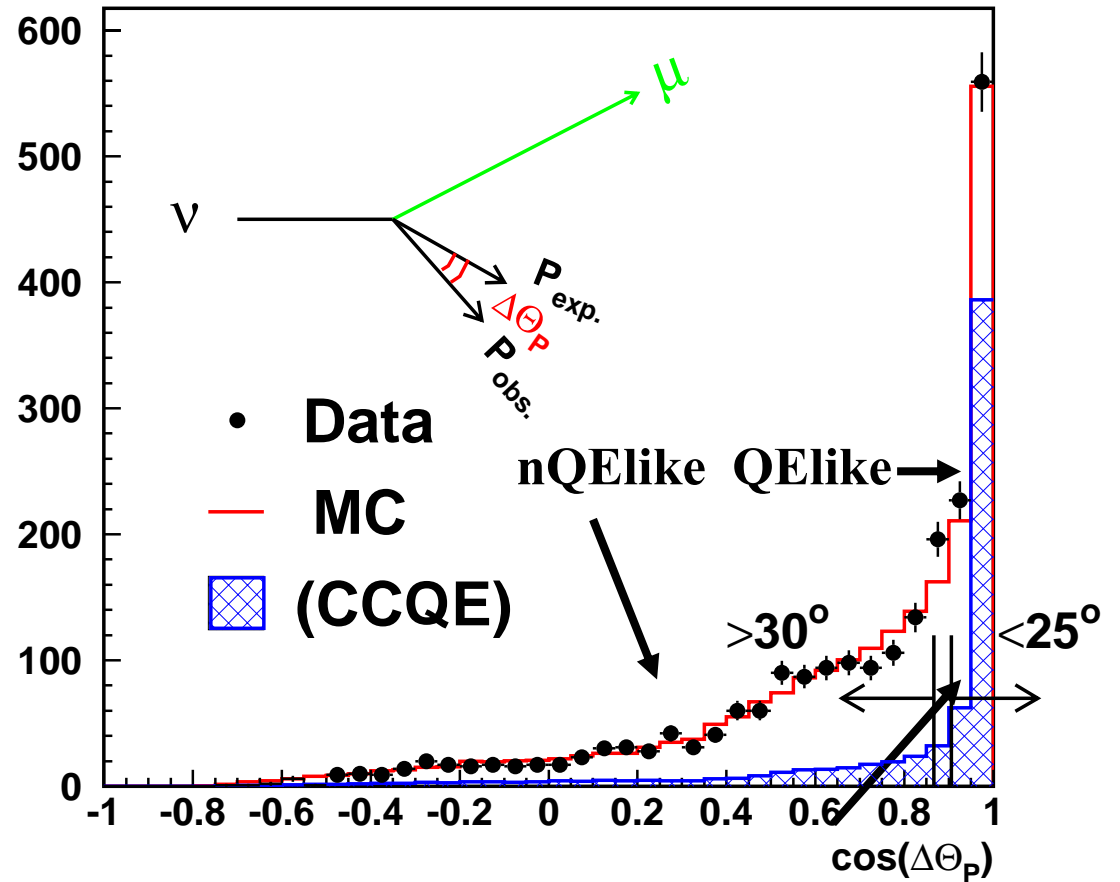
GRV94-Mod.GRV94

Very small effect on oscillation  
analysis

# QE and nQE in SciFi 2track events



## SciFi 2 track $\cos(\Delta\Theta_P)$ distribution



## Used data for $\phi_{\text{near}}(\text{E}\nu)$

### KT

Fully contained in fiducial volume (FCFV)

- No. of events ( $E_{\text{vis}} > 100 \text{ MeV}$ )
- (1) Single ring  $\mu$ -like events

### SciFi

- (2) 1-track  $\mu$  events
- (3) 2-track QE-like events
- (4) 2-track nonQE-like events

4 sets of  $(p_{\mu}, \theta_{\mu})$  distributions

### Pion monitor

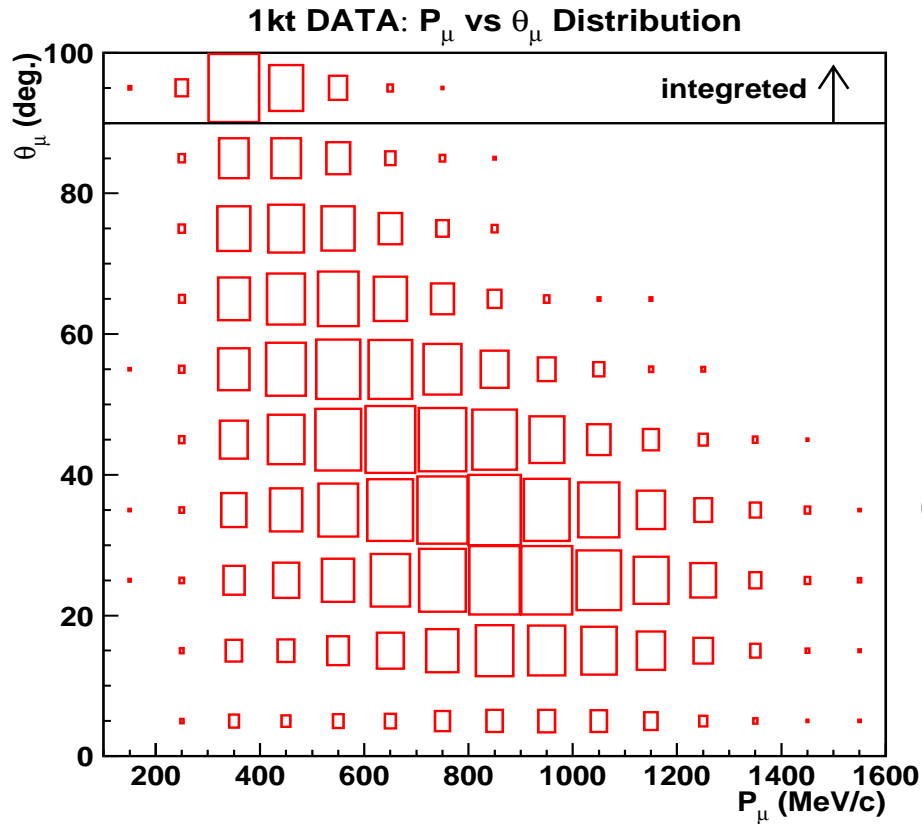
$\pi$  distribution in  $(p_{\pi}, \theta_{\pi}) \rightarrow$  flux estimate  $\phi_{\text{near}}(\text{E}\nu)$

$\nu$  flux  $\phi_{\text{near}}(\text{E}\nu)$  (8 bins)

$\nu$  interaction model (parameterized as **nQE/QE** ratio)

# Fitting method

$$(p_\mu, \theta_\mu) \rightarrow \phi(E\nu), nQE/QE$$



Also  $(p_\mu, \theta_\mu)$  dist. in SciFi

1track, 2track(QE-like), 2track(nQE-like)

$\chi^2=227$  for 197 d.o.f.

Ev

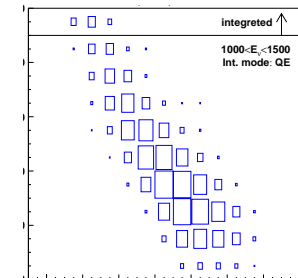
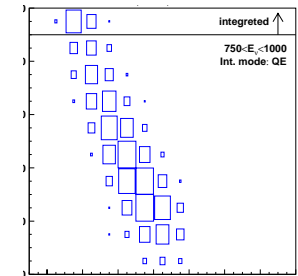
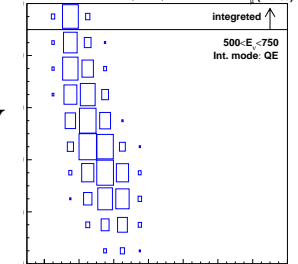
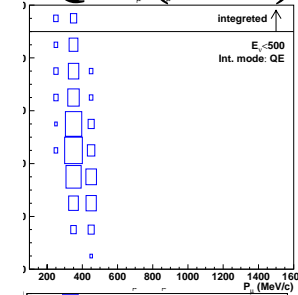
0-0.5 GeV

0.5-0.75GeV

0.75-1.0GeV

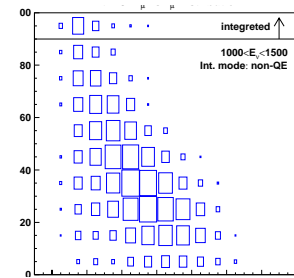
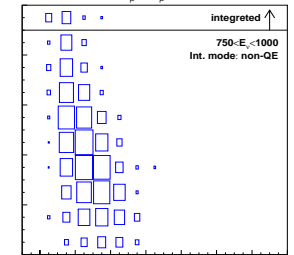
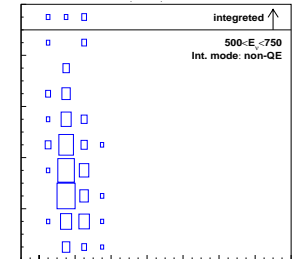
1.0-1.5GeV

QE (MC)



nQE(MC)

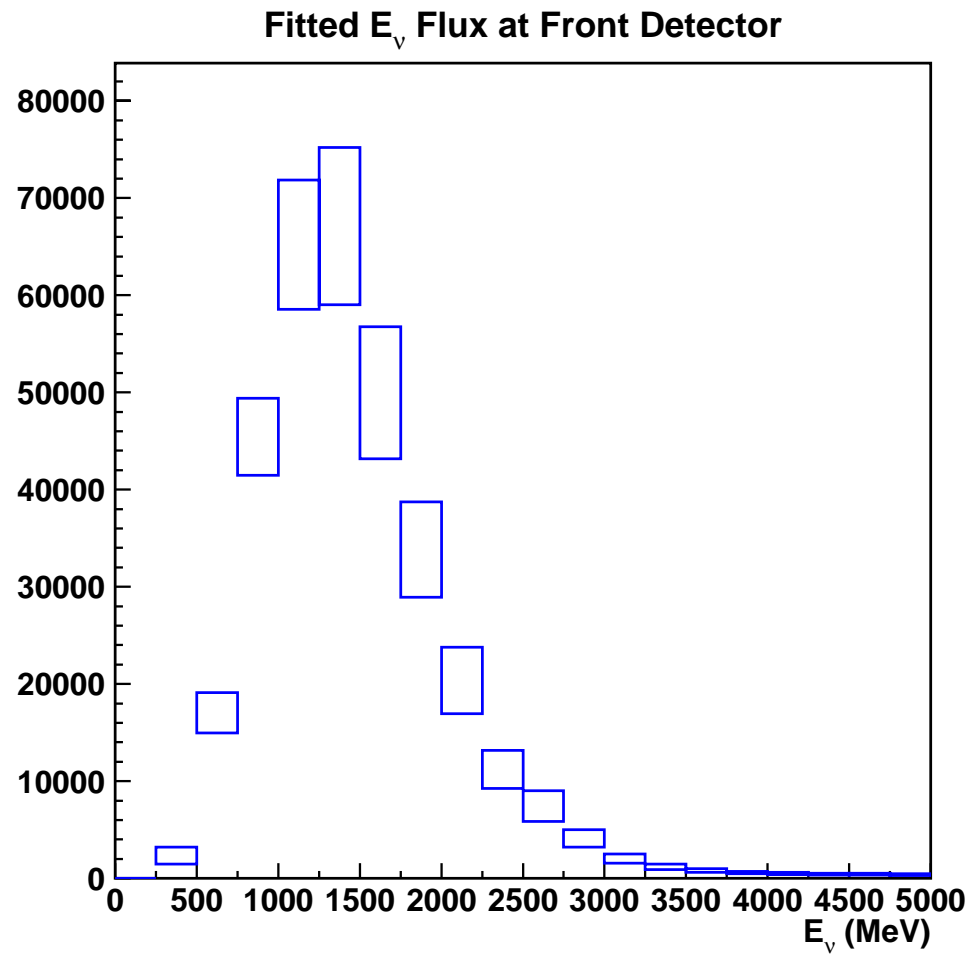
MC templates



•  
•

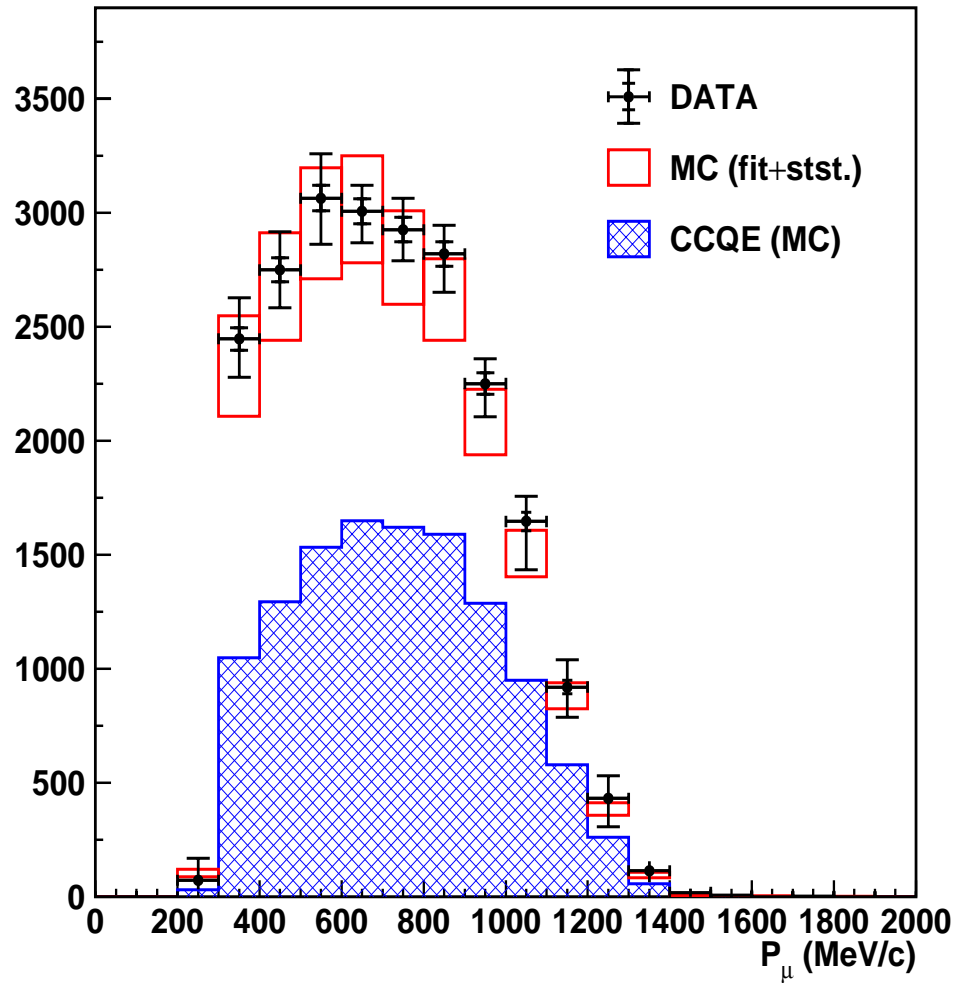
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# Fit result of Neutrino Flux at KEK Site

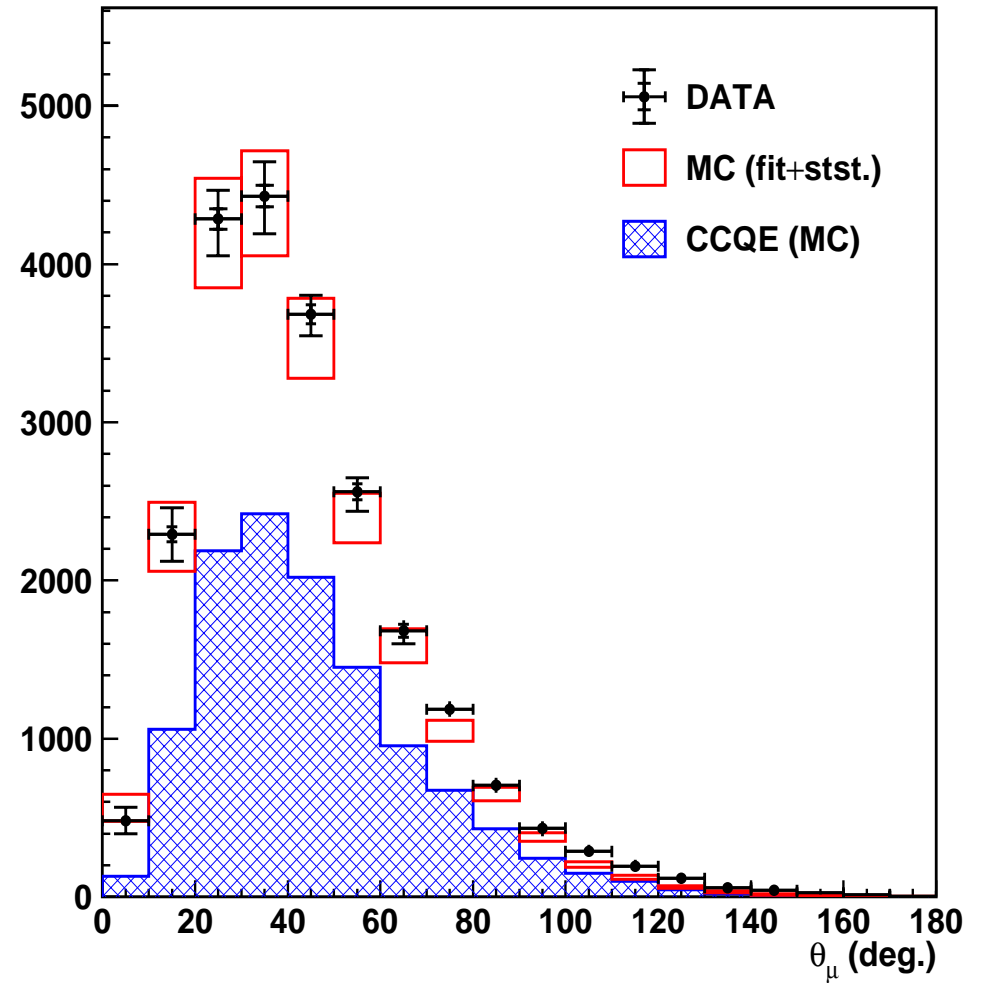


# KT ( $p_\mu, \theta_\mu$ ) distribution using $\phi_{\text{fit}}, \text{QE/nQE}_{\text{fit}}$

1kt:  $\mu$ -momentum Distribution (Fid.25t FC 1-Ring  $\mu$ -like)

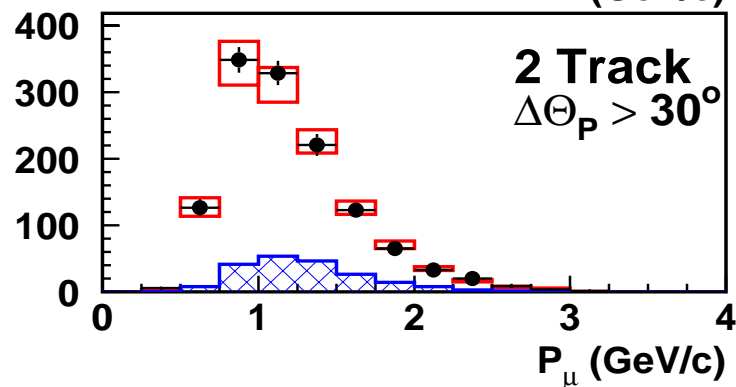
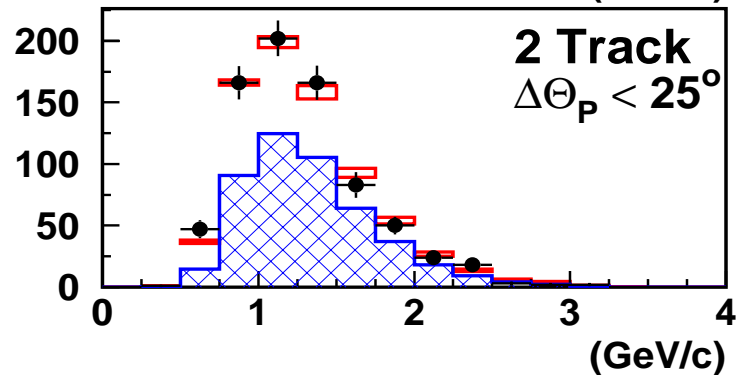
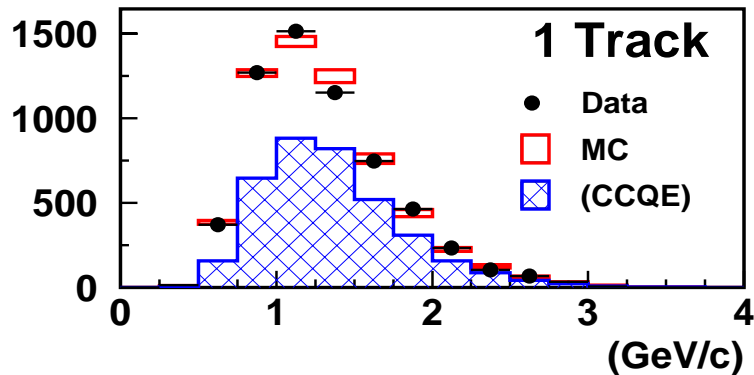


1kt:  $\mu$ -angular Distribution (Fid.25t FC 1-Ring  $\mu$ -like)

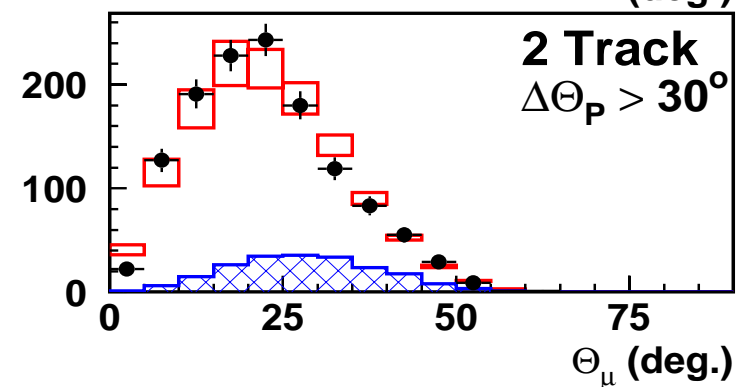
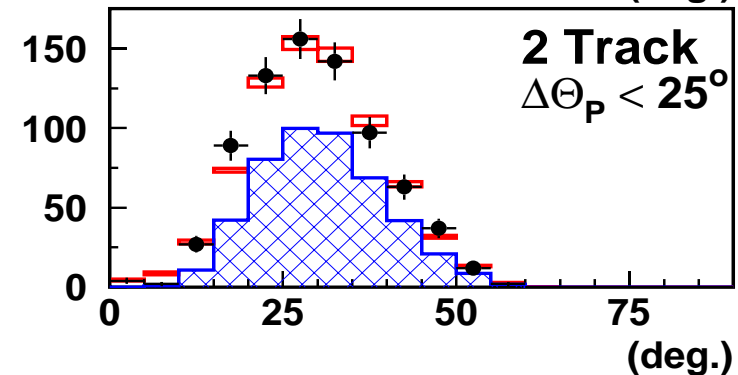
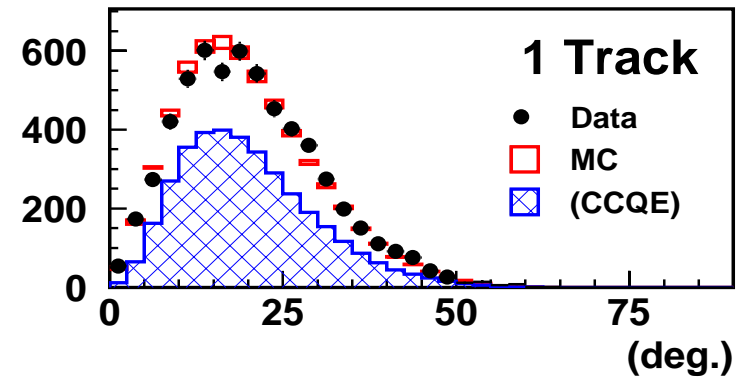


# SciFi $p_\mu, \theta_\mu$ distributions using $\phi_{\text{fit}}, \text{QE/nQE}_{\text{fit}}$

## SciFi $P_\mu$ distributions



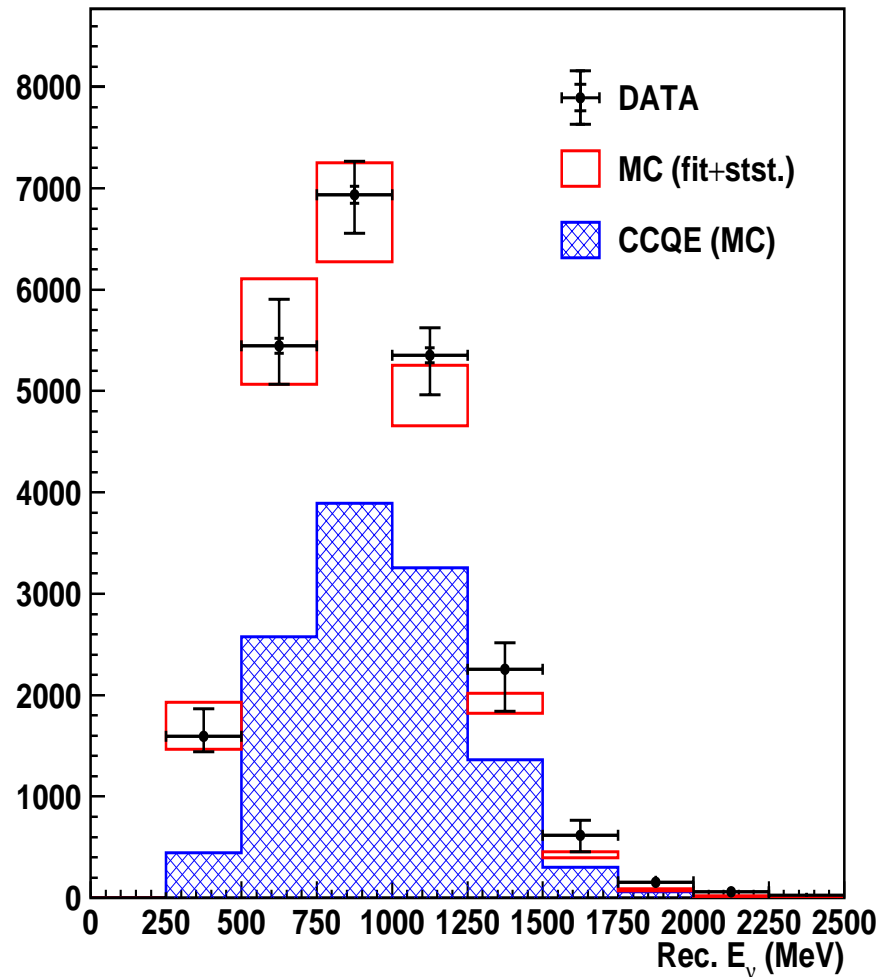
## SciFi $\Theta_\mu$ distributions



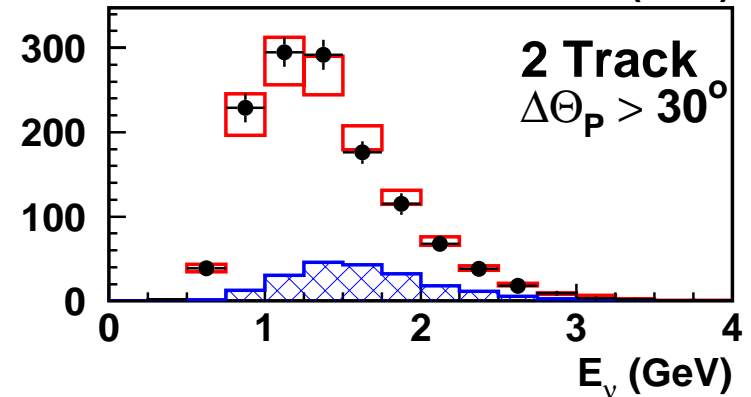
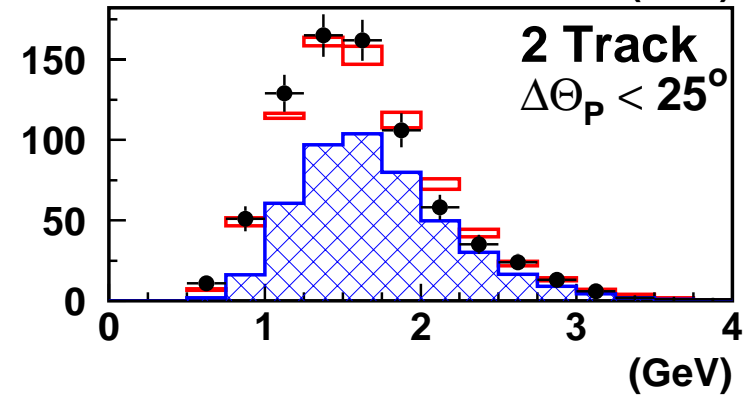
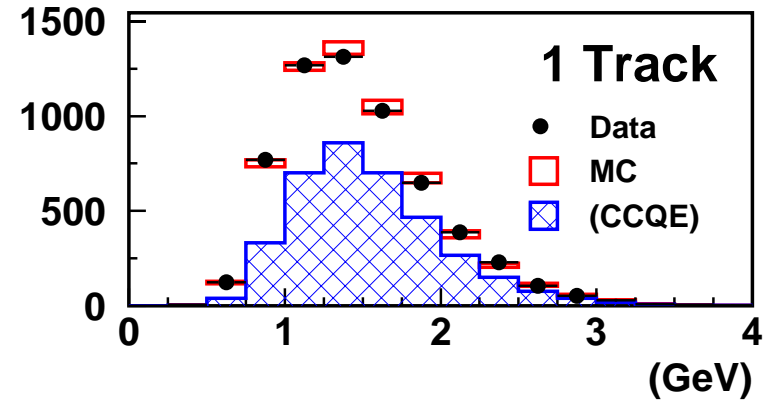


# Reconstructed $E_\nu$ using $\phi_{\text{fit}}$ , $\text{QE/nQE}_{\text{fit}}$

1kt: Reconstructed  $E_\nu$  Distribution (Fid.25t FC 1-Ring  $\mu$ -like)



## SciFi $E_\nu$ distributions



# Far/Near Extrapolation

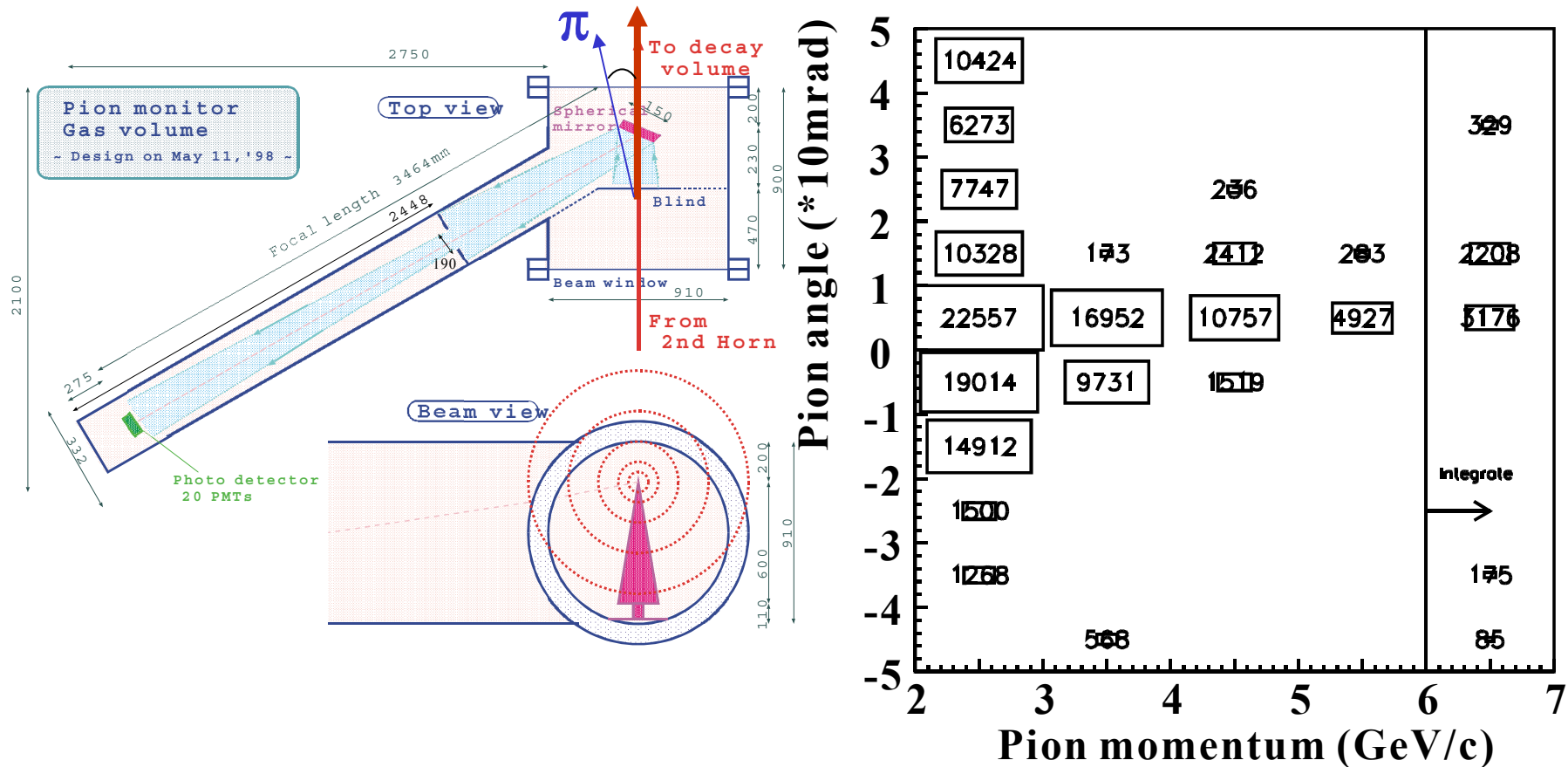


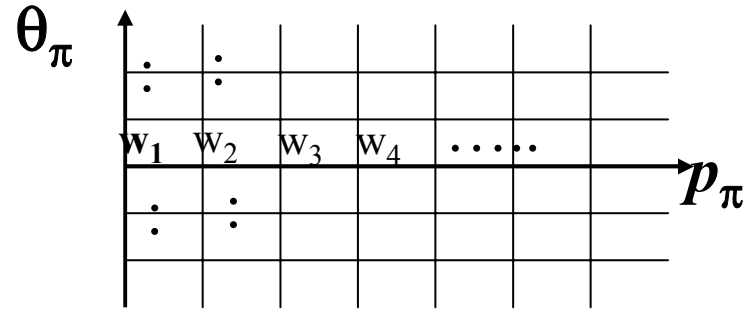
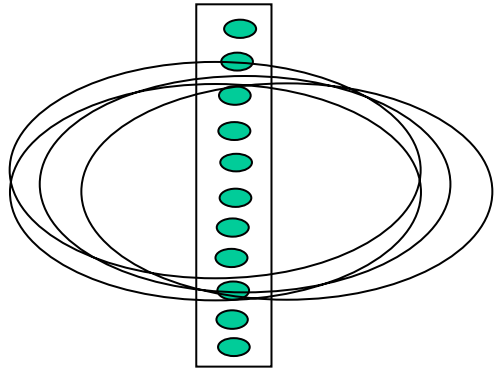
# Pion Monitor

## Far/Near Extrapolation F/N(Ev)

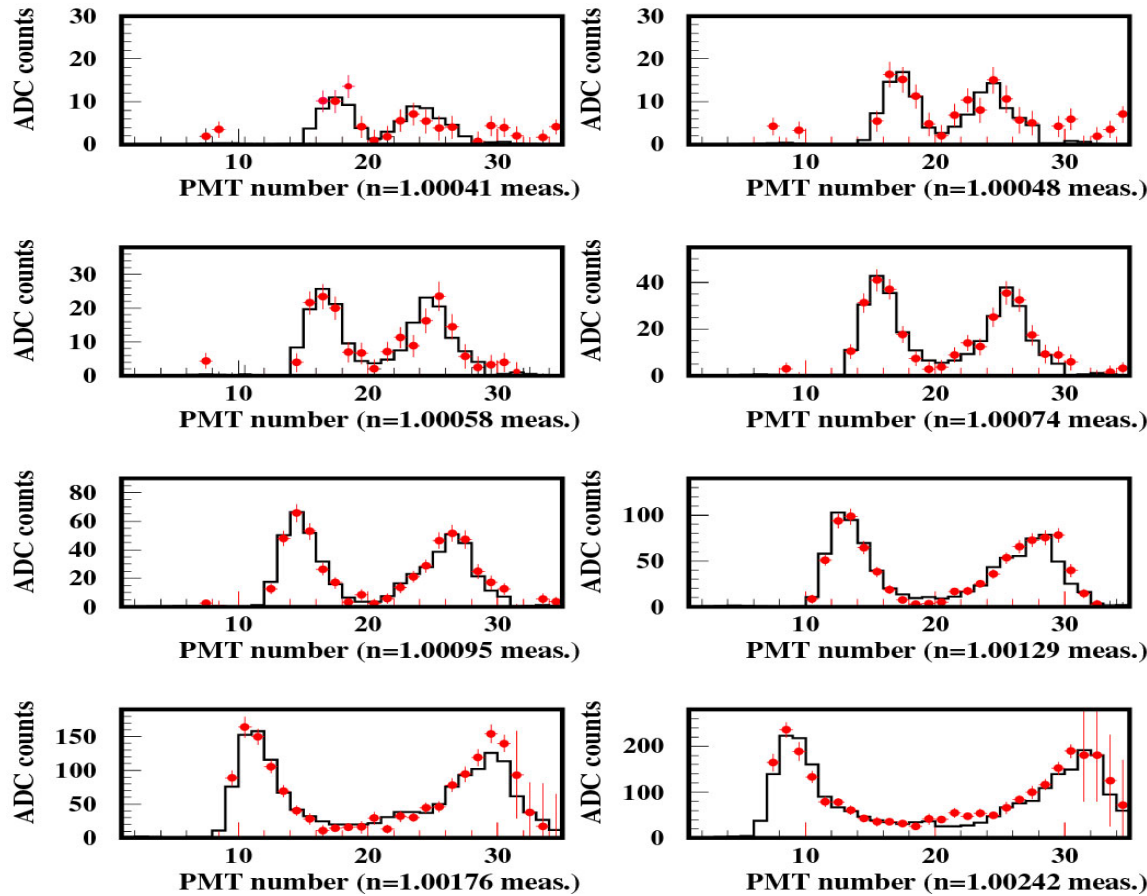
Gas Cherenkov detector: (insensitive to primary protons)

Measure momentum and angular distribution of pions,  $N(p_\pi, \theta_\pi)$  just after the horns.  $p_\pi > 2 \text{ GeV}/c$





**Pion Monitor Fitting (November)**



**Fit with Cherenkov  
light distribution by  
populations in  $(p_\pi, \theta_\pi)$**

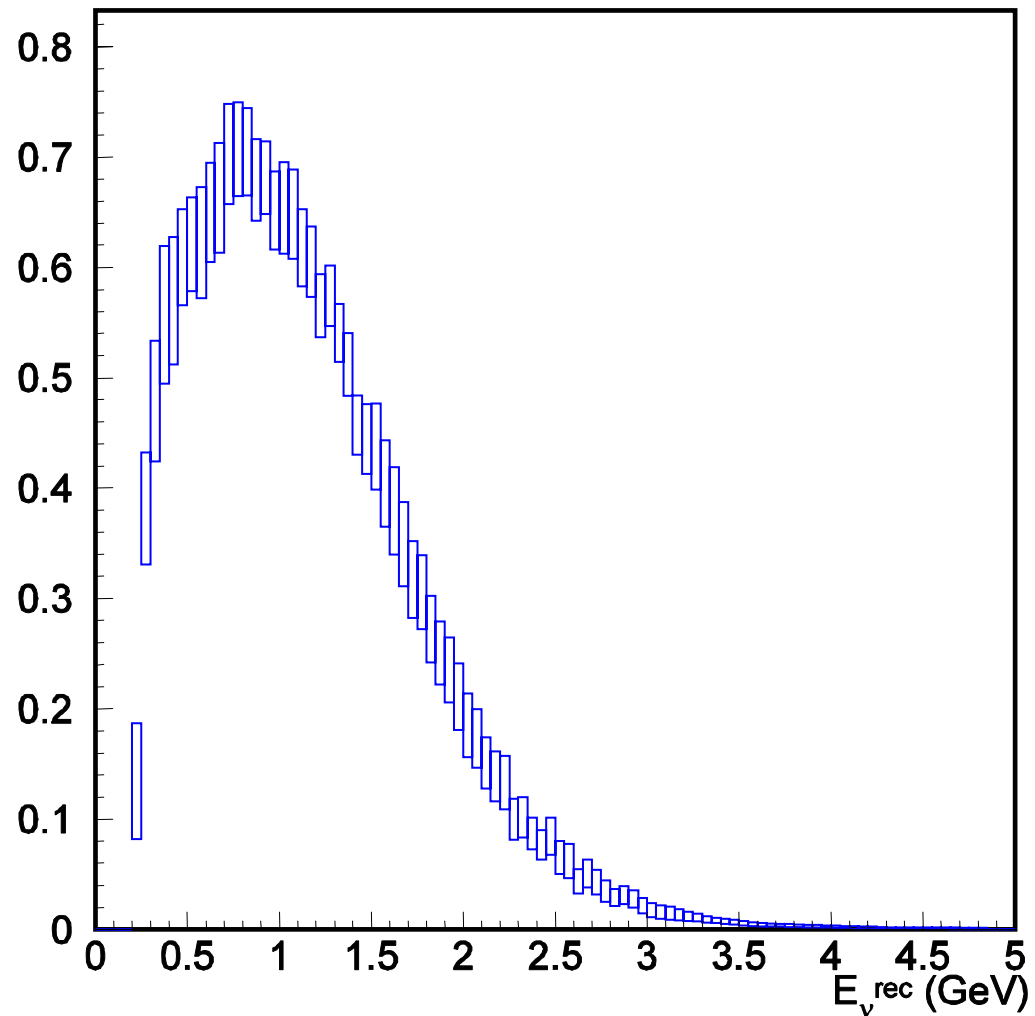
**Error on F/N(Ev)**

**Pion production  
model**

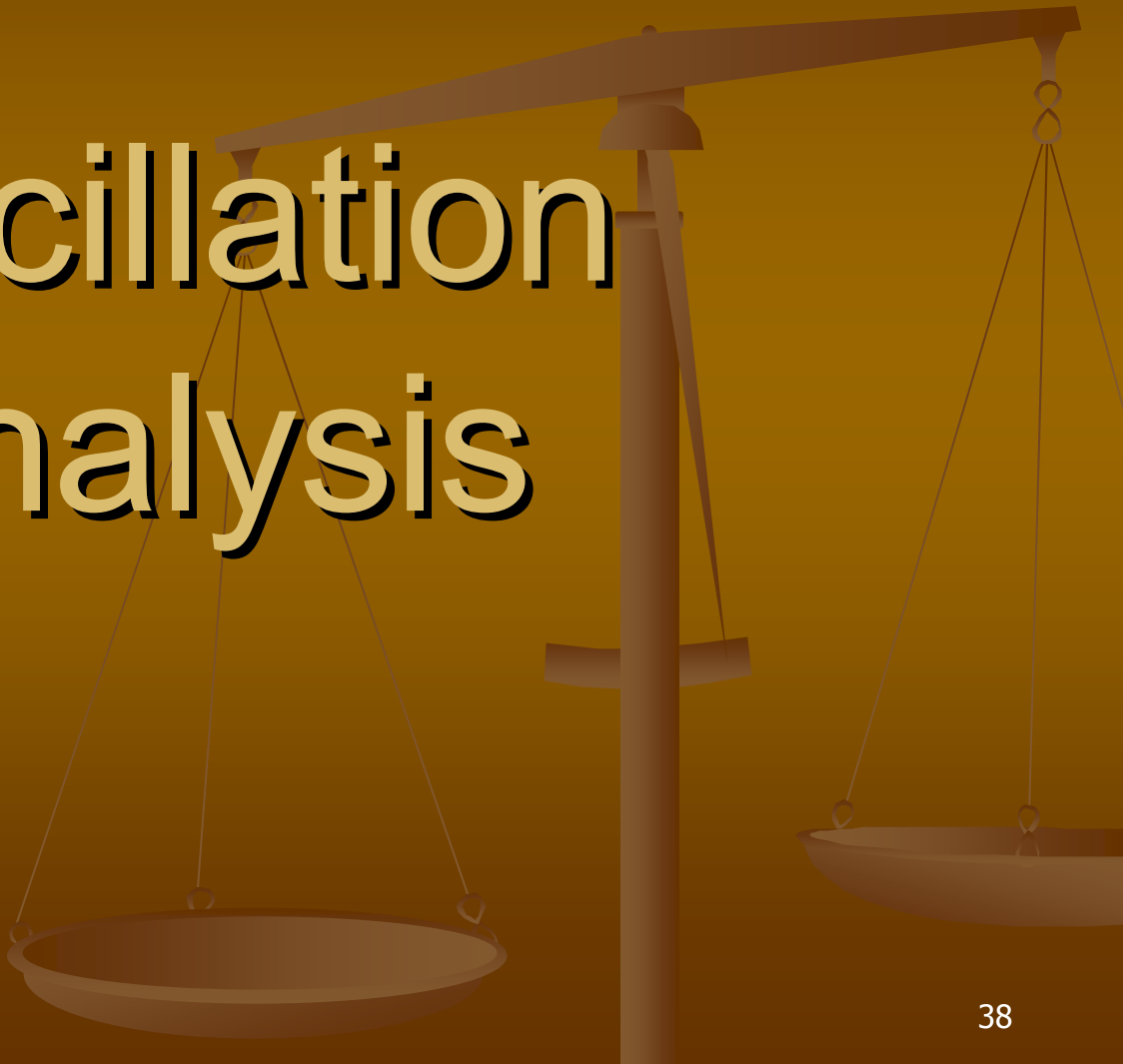
**Beam MC**

# Expected $E_\nu^{\text{rec}}$ spectrum for $1R\mu$

Initial  $1R\mu$  spectrum w/ all syst. err. incl. Escale



# Oscillation Analysis



# Oscillation analysis

Neutrino flux @SK  $\otimes$  *Int. Model*  $\otimes$  Oscillation ( $\sin^2 2\theta, \Delta m^2$ )

## Prediction

- $N_{\text{SK}}(\text{exp}'t)$  : Expected no. of SK events
- $S_{\text{SK}}(E_{\nu}^{\text{rec}})$  :  $1\text{R}\mu$   $E_{\text{rec}}$  distribution(shape)

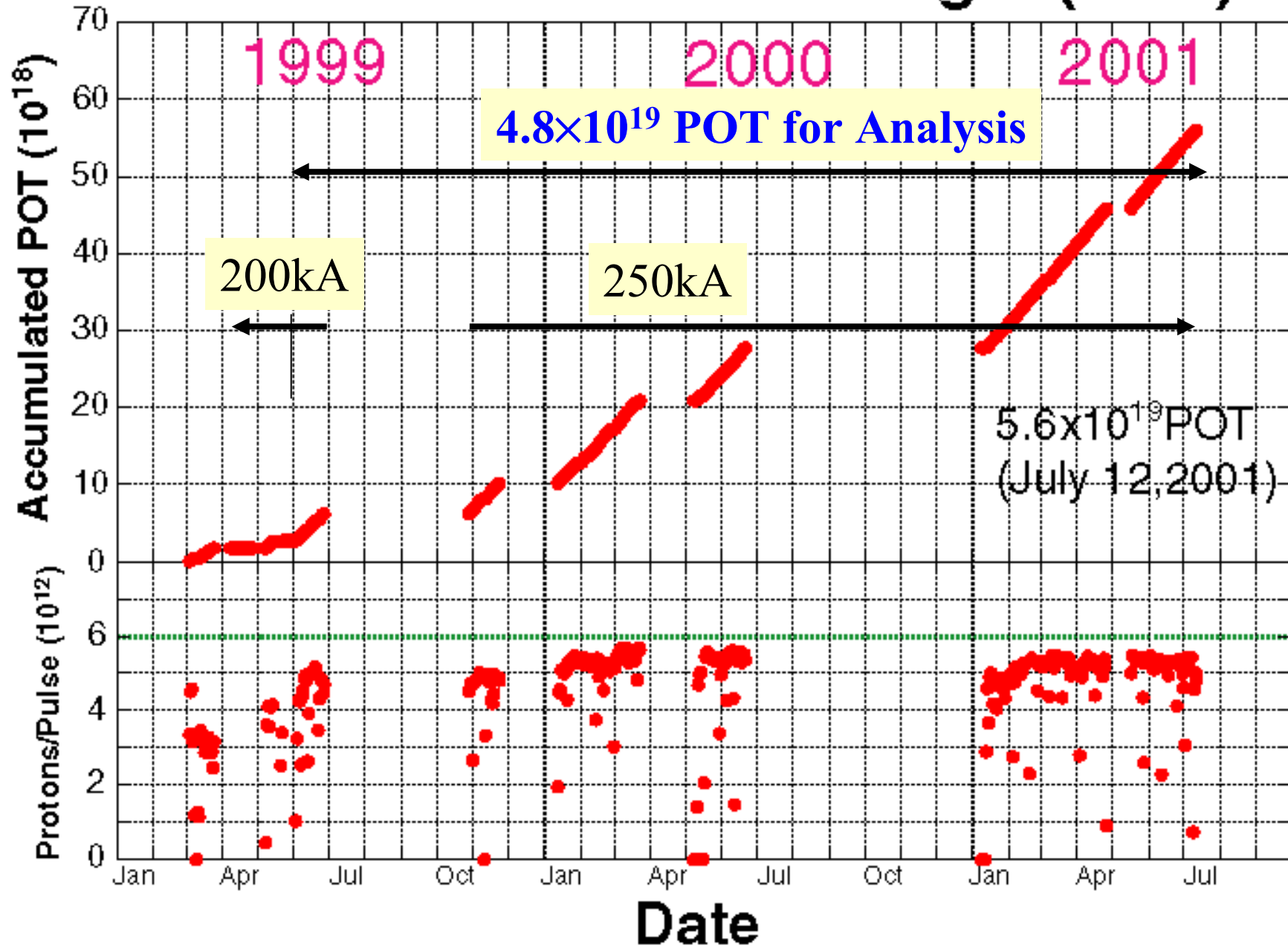
## SK observation

- Observed no. of events in FCFV  $N_{\text{SK}}(\text{obs}, >30\text{MeV})$
- $1\text{R}\mu$  events  $E_{\nu}^{\text{rec}}$  spectrum shape

## Maximum Likelihood Fit in ( $\sin^2 2\theta, \Delta m^2$ )

1. Rejection of Null oscillation hypothesis
2. Contour of allowed region
  - Number of events observed/expected
  - Obs./exp. neutrino energy spectrum shape

# Delivered Protons on Target (POT)





# Data set

- **Data sets**
  - **June 99-July 01 FCFV , Evis>30 MeV**
    - total number of events
    - 56 events observed
  - **Nov 99-July 01 1Rμ events**
    - $E_{\nu}^{\text{rec}}$  shape
    - 29 events observed
- **Running condition**
  - **June 99**
    - Target=2 cm $\phi$  Horn current=200kA (~6.5% of POT)
    - Larger systematic errors in ‘near’ measurements
  - **Nov 99~July 01**
    - Target=3cm $\phi$  Horn current=250kA
    - Full analysis of systematic errors

# Systematic parameters

$$f = (f_{\Phi}, f_{nQE}, f_{F/N}, f_{\varepsilon sk}, f_{Esk}, f_{n6}, f_{n11})$$

$f_{\phi}$  : Flux ( 8 energy bins)

$f_{nQE}$  : QE/nQE ratio

$f_{F/N}$  : Far/Near ratio

$f_{\varepsilon SK}$  : SK reconstruction (Fid, PID, Nring)

$f_{ESK}$  : SK energy scale

$f_{n6}$  : Norm. for June 99

$f_{n11}$  : Norm. Nov 99 ~ Jul 01

# Likelihood

$$L_{tot} = L_{norm}(f) \cdot L_{shape}(f) \cdot L_{syst}(f)$$

## Normalization term

$$L_{norm} = Poisson(N_{obs}, N_{exp}(f))$$

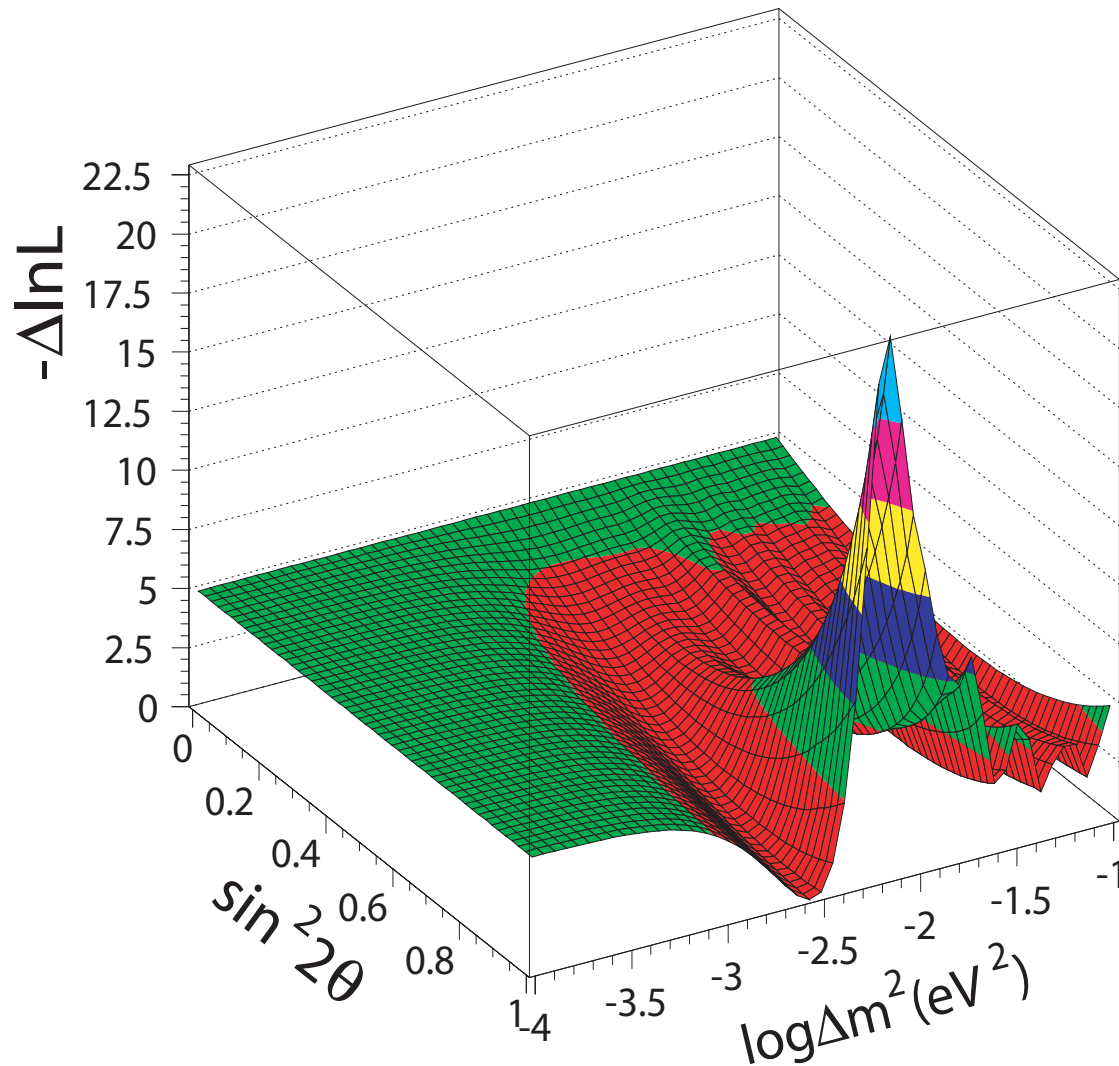
## Shape term for FCFV 1Rμ

$$L_{shape} \equiv \prod_{i=1}^{29} P((f_{Esk} \cdot E_i), \Delta m^2, \sin^2 2\theta, f)$$

## Systematic parameter constraint term

$$L_{syst} \equiv \exp\left(-\Delta f_{\Phi, nQE}^T \cdot M_{FD}^{-1} \cdot \Delta f_{\Phi, nQE} / 2\right) \bullet \bullet \bullet \bullet \bullet \\ \times \exp\left(-f_{n6}^2 / 2\sigma_{n6}^2\right) \exp\left(-f_{n11}^2 / 2\sigma_{n11}^2\right) \exp\left(-\Delta f_{Esk}^2 / 2\sigma_{Esk}^2\right)$$

# 3d plots of $\Delta\ln L$ for shape+norm



**L at  $(\Delta m^2, \sin^2 2\theta)$**

- **analysis-1**

**Maximize L by  
adjusting systematic  
parameters.**

- **analysis-2**

**The MC generation of  
the systematic  
parameters &  
L=the mean values.**

# Null Oscillation Probability

## Null Oscillation Probability

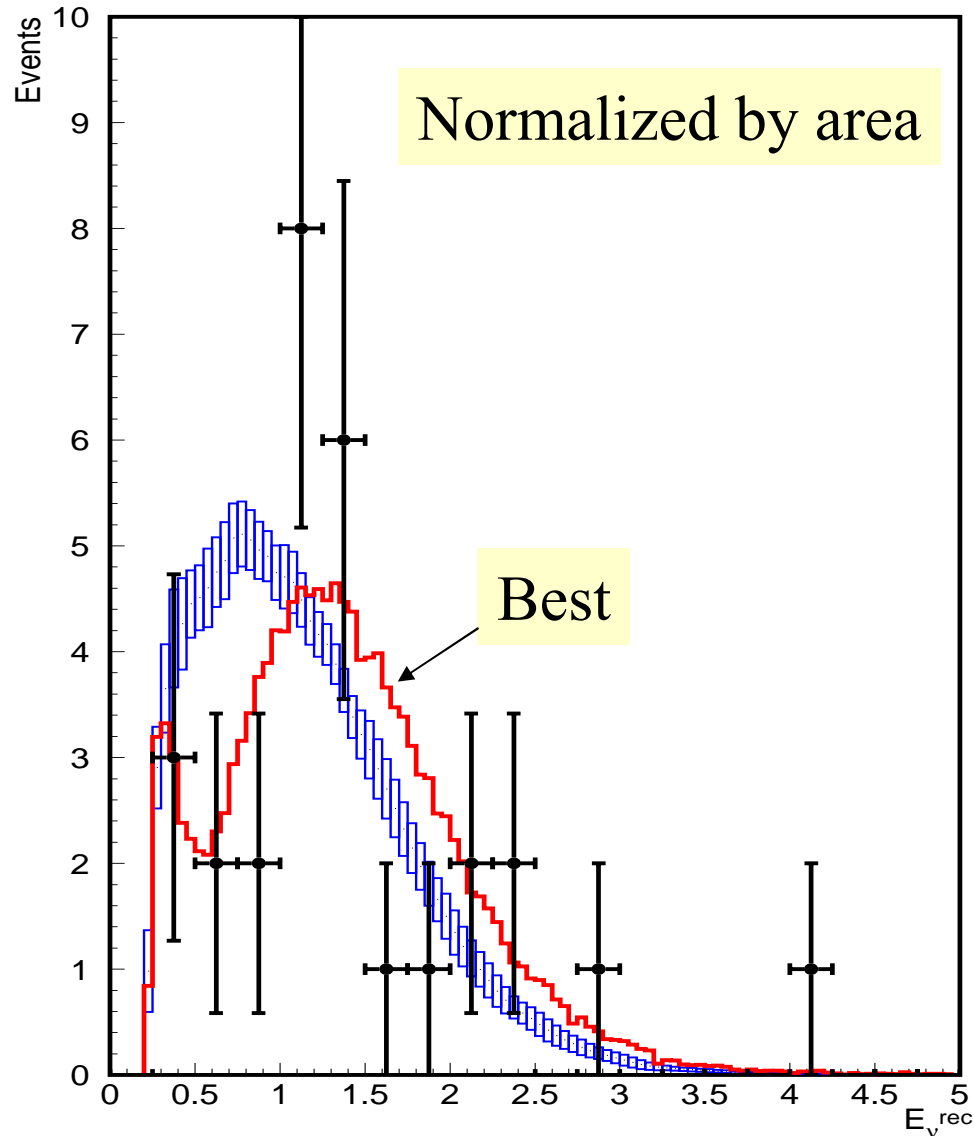
	analysis-1	analysis-2
$N_{SK}$ only	1.3%	0.7%
Shape only	15.7%	14.3%
$N_{SK}$ +Shape	0.7%	0.4%

## Best fit ( $\sin^2 2\theta$ , $\Delta m^2$ )

Shape only	(1.0, $3.0 \times 10^{-3} \text{eV}^2$ )	(1.0, $3.2 \times 10^{-3} \text{eV}^2$ )
(Allowing unphys.)	(1.09, $3.0 \times 10^{-3} \text{eV}^2$ )	(1.05, $3.2 \times 10^{-3} \text{eV}^2$ )
$N_{SK}$ +Shape	(1.0, $2.8 \times 10^{-3} \text{eV}^2$ )	(1.0, $2.7 \times 10^{-3} \text{eV}^2$ )
(Allowing unphys. )	(1.03, $2.8 \times 10^{-3} \text{eV}^2$ )	(1.05, $2.7 \times 10^{-3} \text{eV}^2$ )

**Both Shape and  $N_{SK}$  +Shape indicate consistent parameter region**

# Best fit 1R $\mu$ spectrum & N<sub>SK</sub>



**Best fit point ( $\sin^2 2\theta$ ,  $\Delta m^2$ )**  
**= (1.0,  $2.8 \times 10^{-3} \text{eV}^2$ )**

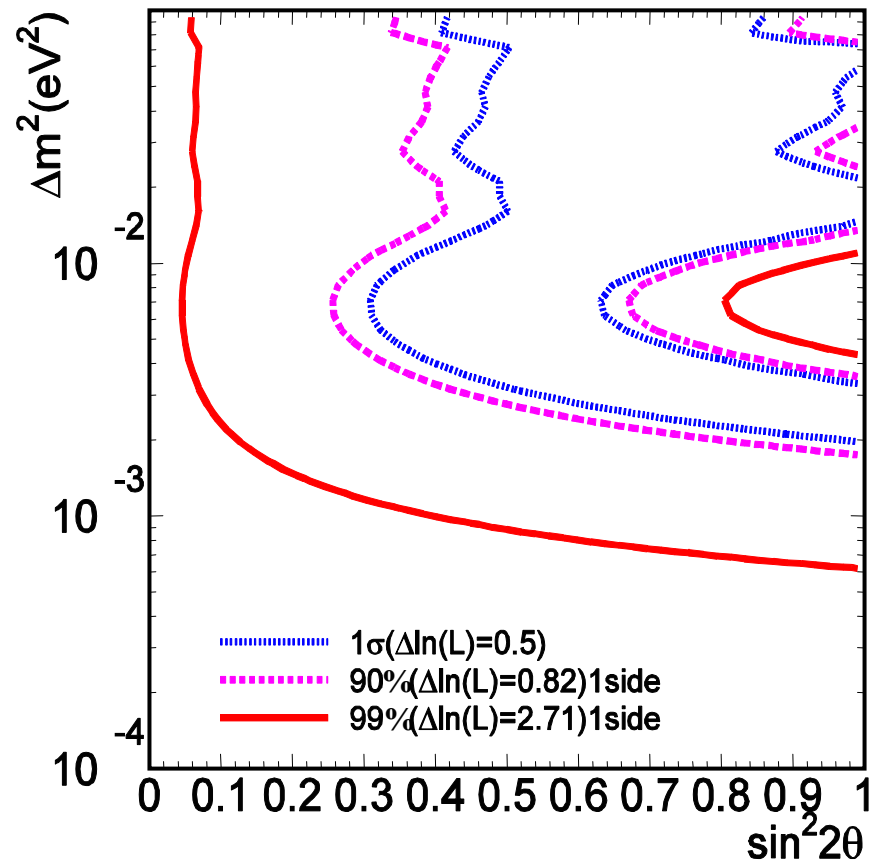
KS test prob.(shape): 79%

**$N_{SK} = 54$  (Obs. = 56)**

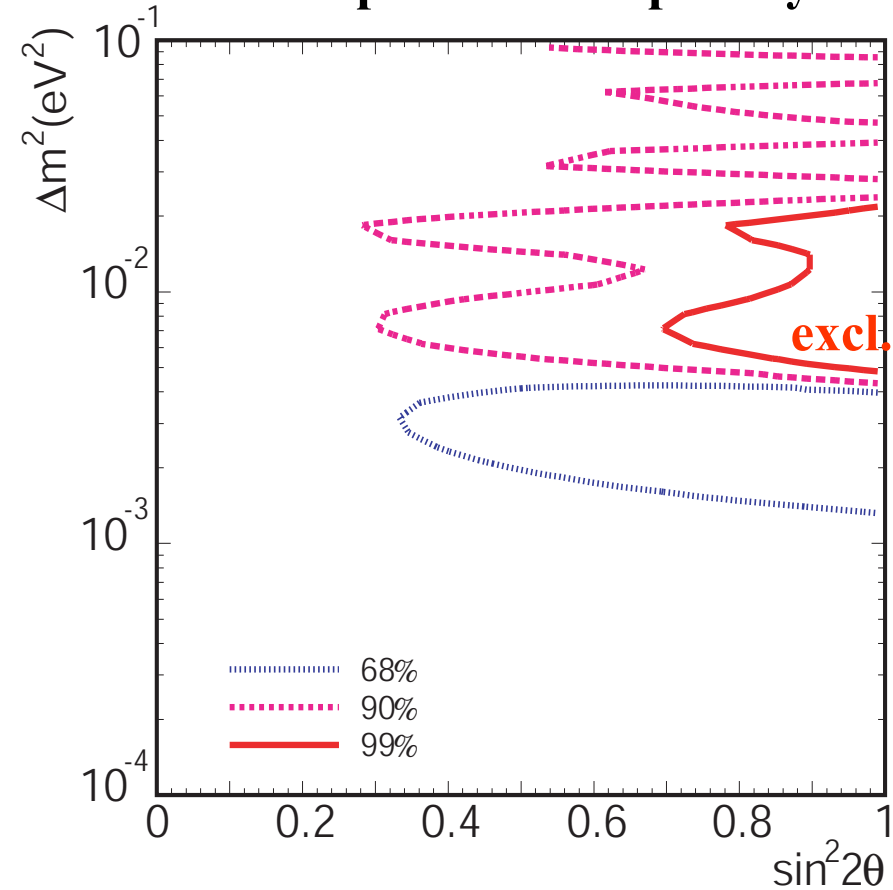
**Very good agreement**  
**Shape &  $N_{SK}$**

# Allowed regions

Total no. of Events only

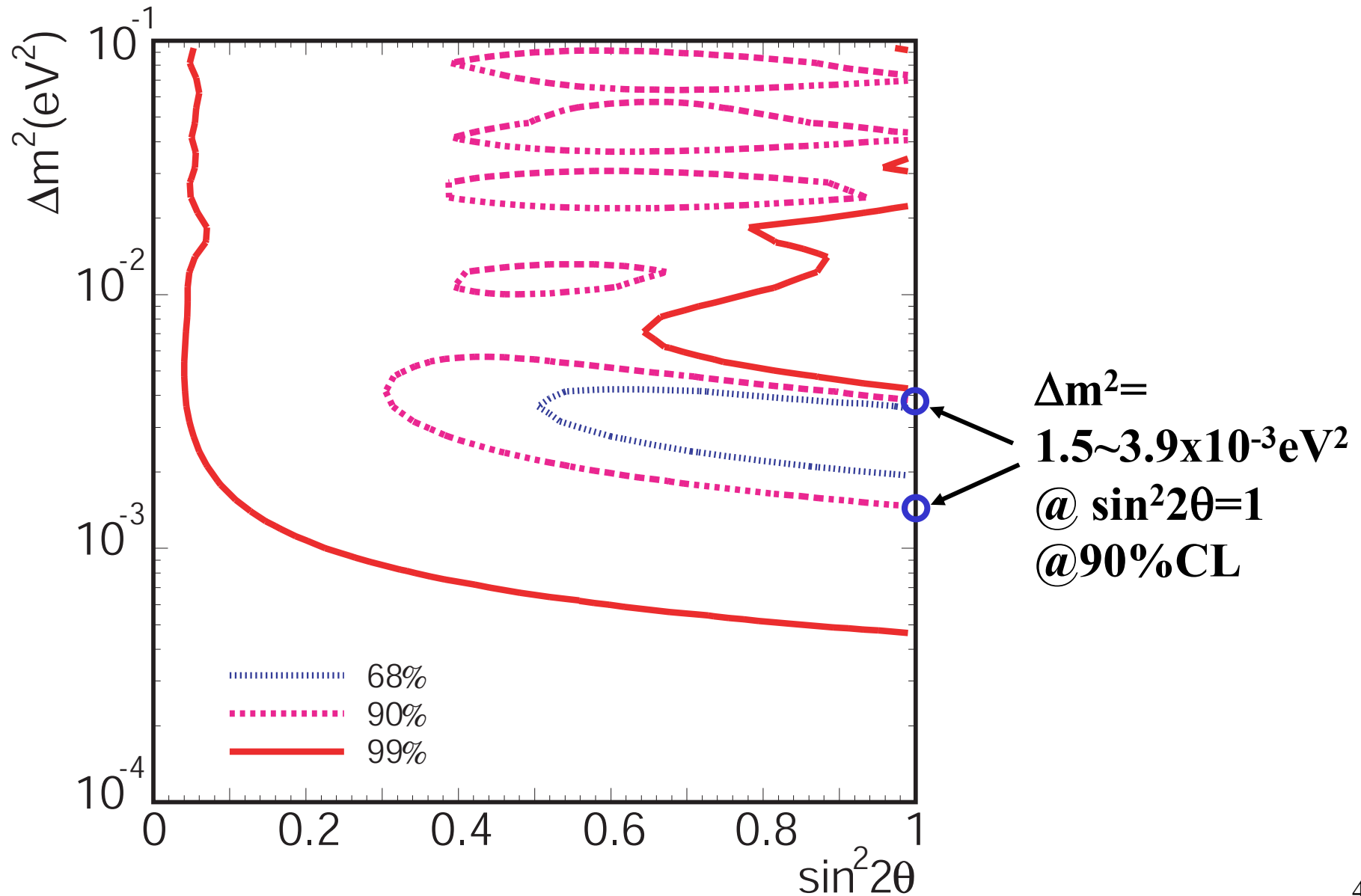


Spectrum Shape only



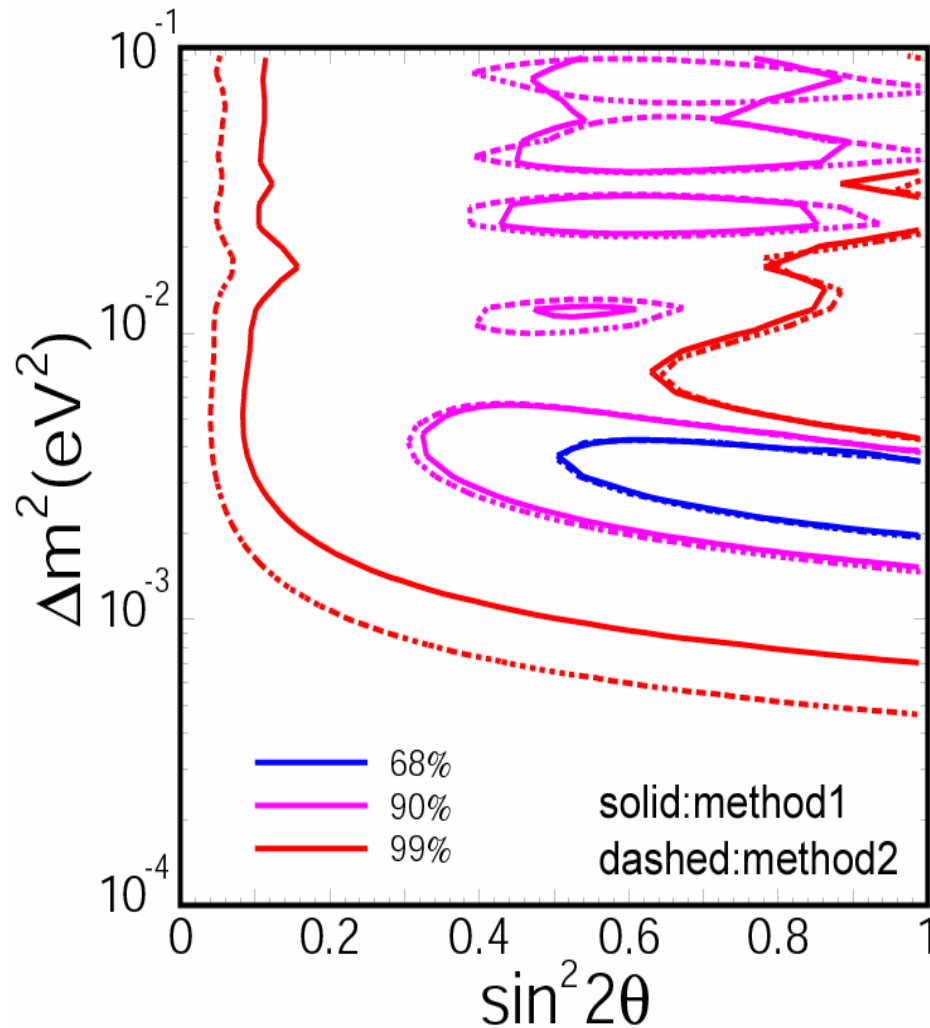
**Both indicate consistent  $\Delta m^2$  region**

# Allowed region (Shape+Norm)

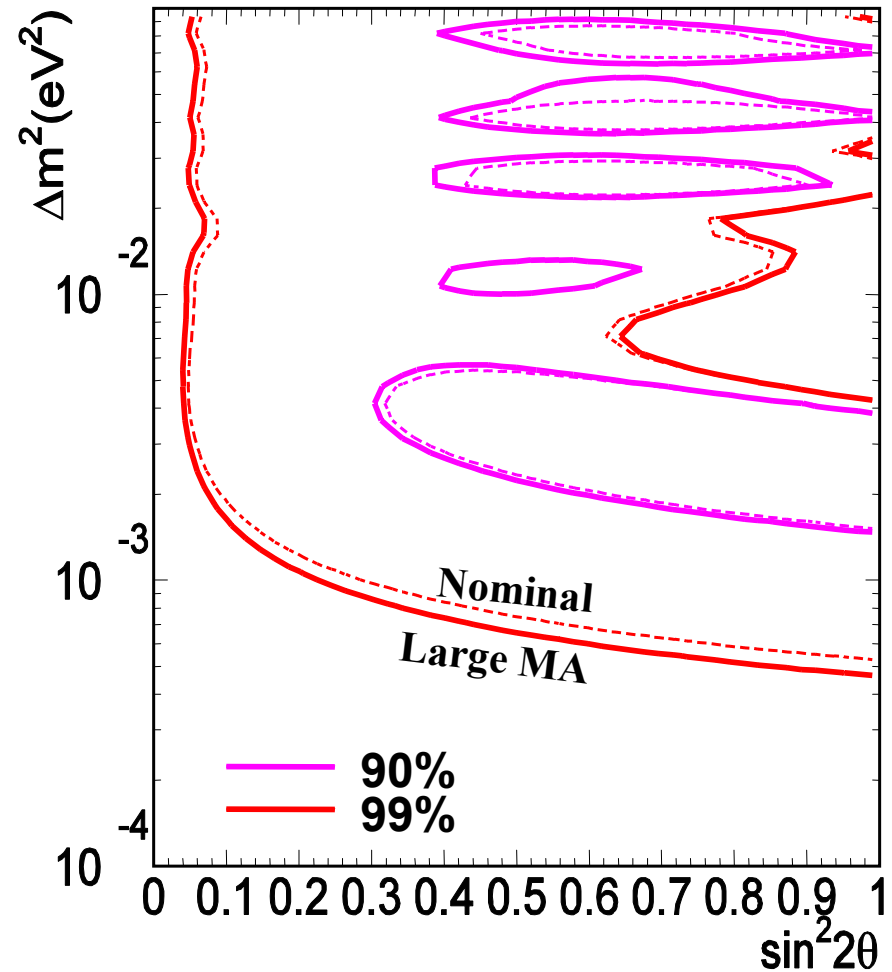




# Comparison with diff. L & form factor MA



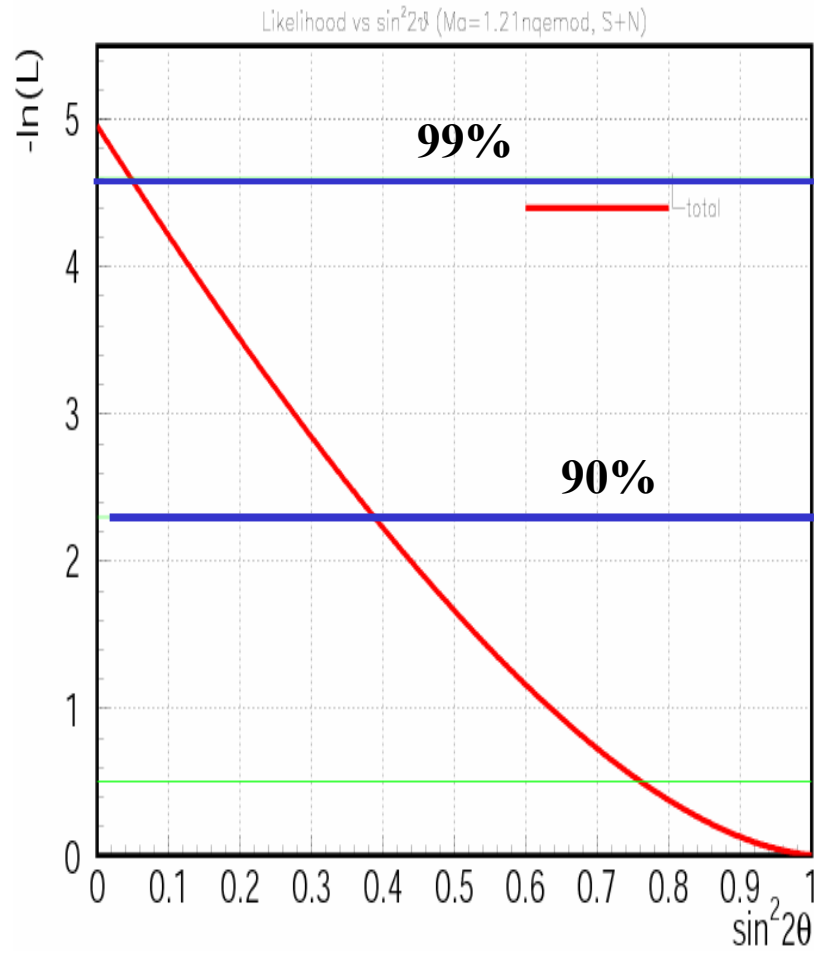
**Reasonable agreement  
btw definition of L**



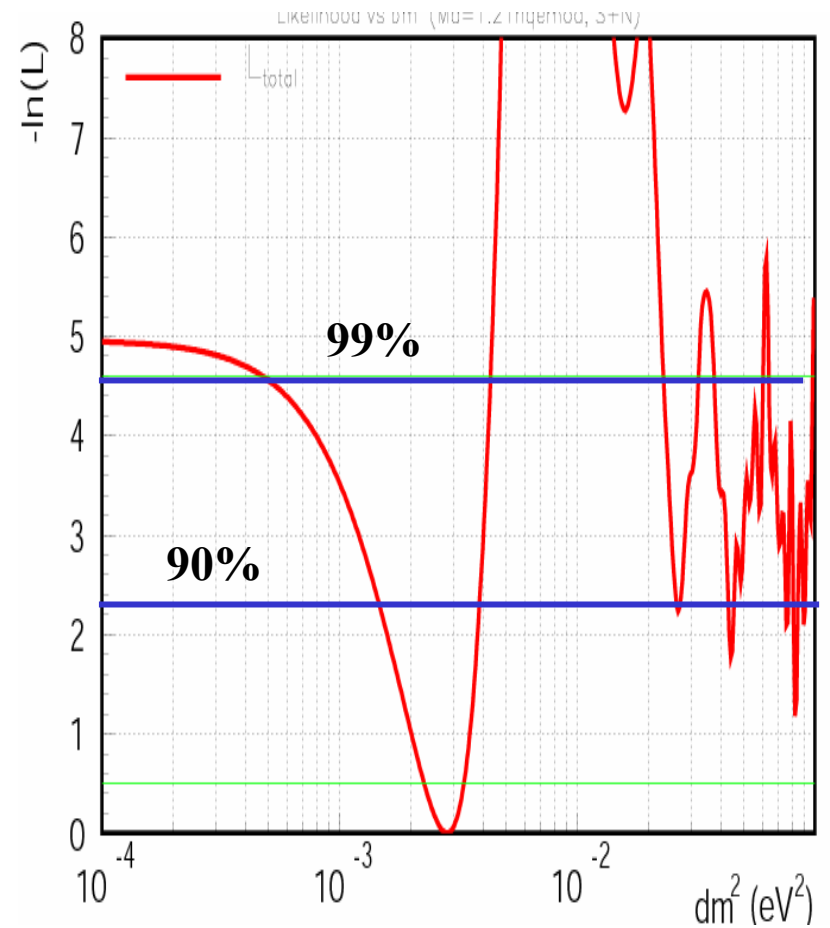
**Change of  $\nu$  interaction  
model has small effect**

# $\sin^2 2\theta$ and $\Delta m^2$

2002/05/14 19:56



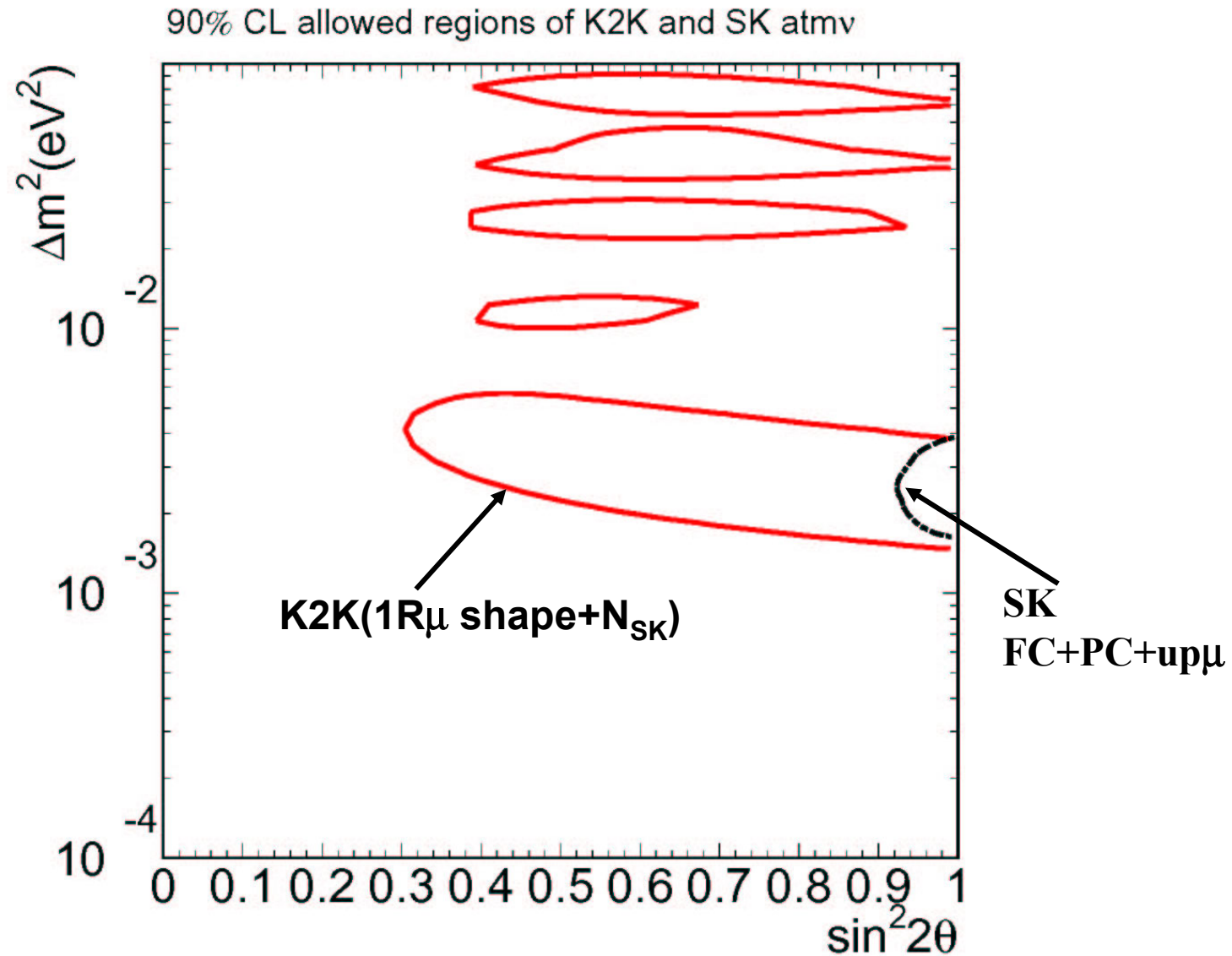
$\sin^2 2\theta$



$\Delta m^2$

# Comparison with SK atm $\nu$ observation

## 90% allowed region



# Conclusion

- **K2K Oscillation analysis on June99 ~July01 data**
  - Full error analysis

1. **Null oscillation probability is less than 1%**
2. **Both SK rate reduction and  $E_\nu^{\text{rec}}$  shape indicate consistent oscillation parameters region**
3.  **$\Delta m^2=1.5\sim 3.9\times 10^{-3}\text{eV}^2$  for  $\sin^2 2\theta=1$  @ 90%CL**
4.  **$\sin^2 2\theta, \Delta m^2$  are consistent with atmospheric neutrino results**

- The best fit point ( $\sin^2 2\theta=1.0, \Delta m^2=2.8\times 10^{-3}\text{ eV}^2$ )

cf. Atmospheric neutrino results

$\Delta m^2=(1.6\sim 3.9)\times 10^{-3}\text{ eV}^2$  for  $\sin^2 2\theta=1.0$

best fit ( $\sin^2 2\theta=1.0, \Delta m^2=2.5\times 10^{-3}\text{ eV}^2$ )

- **Data taking will resume within this year**