# Toward discovery of $\theta_{13}$ and the CPV

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# Introduction

#### Evidences of v oscillation

- Atm v (SK,1998), confirmed by K2K (2004)
- Solar v (SK+SNO,2001), confirmed by KamLAND (2002, 2004)
- Finite masses!! & Large mixings!!!!! Very much different from quarks
- First evidence beyond SM

#### Next steps

- Understand whole structure of  $\nu$  mass/mixing
- How similar/different from quark sector?
  - "Standard" mixing w/ 3x3 matrix?
  - Mass hierarchy?
  - CP violation?

Would lead physics beyond SM

- Next generation LBL experiments and reactor exp.
  - w/ High statistics and small systematics

#### New era of precision "Neutrino Flavor Physics"

- Cf. Have been done last ~40yrs for quark sector

# 3 flavor mixing

$$\left| \mathcal{V}_{l} \right\rangle = \Sigma U_{li} \left| \mathcal{V}_{i} \right\rangle$$
 *m<sub>i</sub>*: 3 masses,  
Weak Mass eigenstates

Maki-Nakagawa-Sakata Matrix  $(s_{ij}=\sin\theta_{ij}, c_{ij}=\cos\theta_{ij})$ 

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} = \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \cdot \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \cdot \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & e^{-i\delta} \end{pmatrix} \cdot \begin{pmatrix} c_{13} & 0 & s_{13} \\ 0 & 1 & 0 \\ -s_{13} & 0 & c_{13} \end{pmatrix}$$

Parameters governing oscillation

- 3 mixing angles ( $\theta_{12}$ ,  $\theta_{23}$ ,  $\theta_{13}$ )
- 1 CPV phase ( $\delta$ )
- 2 (indep.) mass differences  $(\Delta m_{ij} = m_i^2 m_j^2)$

**Oscillation probabilities**  
when 
$$\begin{cases} \Delta m_{12}^{2} \ll \Delta m_{23}^{2} \approx \Delta m_{13}^{2} \\ E_{\nu/L} \approx \Delta m_{23}^{2} \end{cases}$$
 contribution from  $\Delta m_{12}$  is small  
(No CPV & matter eff. approx.)  
 $\nu_{\mu}$  disappearance (LBL/Atm)  $\implies \theta_{23}$  and  $\Delta m_{23}^{2}$   
 $P_{\mu \to x} \approx 1 - \cos^{4} \theta_{13} \cdot \sin^{2} 2\theta_{23} \cdot \sin^{2} (1.27\Delta m_{23}^{2}L/E_{\nu})$   
 $\nu_{e}$  appearance (LBL/Atm)  $\implies \theta_{13}$  and  $\Delta m_{13}^{2}$   
 $P_{\mu \to e} \approx \sin^{2} \theta_{23} \cdot \sin^{2} 2\theta_{13} \cdot \sin^{2} (1.27\Delta m_{13}^{2}L/E_{\nu})$   
 $\nu_{e}$  disappearance (Reactor)  $\implies$  Pure  $\theta_{13}$  and  $\Delta m_{13}^{2}$   
 $P_{e \to x} \approx 1 - \sin^{2} 2\theta_{13} \cdot \sin^{2} (1.27\Delta m_{13}^{2}L/E_{\nu})$   
 $\sum_{k=0.5}^{k} e^{k} - \sin^{2} 2\theta_{13} \cdot \sin^{2} (1.27\Delta m_{13}^{2}L/E_{\nu})$ 

# **Results from atm & acc.** $v_{\mu}$ disapp.( $\theta_{23}$ , $\Delta m_{23}^2$ )



# **Results from solar(** $v_e$ **) & reactor(** $\overline{v}_e$ **) (** $\theta_{12}$ , $\Delta m_{12}^2$ **)**

- Solar neutrino observations
  - SK, SNO
    - $\nu_e \rightarrow \nu_\mu, \nu_\tau$
- KamLAND
  - Reactor  $\overline{v}_e$  disappearance
  - Few MeV x ~200km

Global (solar+KamLAND) Best fit  $sin^2\theta_{sol} = 0.29 (\theta_{sol}=32.6^{\circ})$  $\Delta m_{sol}^2 = 8.1x10^{-5}eV^2$ 

3 $\sigma$  region: 0.23< $\sin^2\theta_{sol}$ <0.37 (28.7°< $\theta_{sol}$ <37.5°)  $\Delta m_{sol}^2$ =7.3~9.1x10<sup>-5</sup>eV<sup>2</sup> T.Kobayashi (KEK)



# **Latest constraint on** $\theta_{13}$



# **Present knowledge and What's next?**



- Only unknown mixing  $\theta_{13}$  (and really  $\Delta m_{13}^2 \sim \Delta m_{23}^2$ ?)
- Mass hierarchy (sign of  $\Delta m^2$ )
- CPV
- Approaches
  - LBL experiment: Multi purpose ( $\theta_{13}$ , sign( $\Delta m^2$ ),CPV,  $\theta_{23}$ , $\Delta m_{23}^2$ )
  - Reactor-based  $\overline{\nu}_{e}$  disappearance: single purpose ( $\theta_{13}$ ), complementary

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# $v_{\mu} \rightarrow v_{e}$ appearance and CPV

$$\begin{split} P(\nu_{\mu} \rightarrow \nu_{e}) &= 4C_{13}^{2}S_{23}^{2}\sin^{2}\frac{\Delta m_{31}^{2}L}{4E} \times \left(1 + \frac{2a}{\Delta m_{31}^{2}}\left(1 - 2S_{13}^{2}\right)\right) \quad \text{Main} \\ &+ 8C_{13}^{2}S_{12}S_{13}S_{23}(C_{12}C_{23}\cos\delta - S_{12}S_{13}S_{23})\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\sin\frac{\Delta m_{21}^{2}L}{4E} \\ &- 8C_{13}^{2}C_{12}C_{23}S_{12}S_{13}S_{23}\sin\delta\sin\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\sin\frac{\Delta m_{21}^{2}L}{4E} \quad \text{CP-odd} \\ &+ 4S_{12}^{2}C_{13}^{2}\left\{C_{12}^{2}C_{23}^{2} + S_{12}^{2}S_{23}^{2}S_{13}^{2} - 2C_{12}C_{23}S_{12}S_{23}S_{13}\cos\delta\right\}\sin^{2}\frac{\Delta m_{21}^{2}L}{4E} \quad \text{Solar} \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}^{2}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \quad \text{Matter} \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}^{2}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \quad \text{Matter} \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}^{2}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \quad \text{Matter} \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \quad \text{Matter} \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \quad \text{Matter} \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \quad \text{Matter} \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}^{2}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}^{2}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \\ &- 8C_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^{2}S_{13}^$$

$$A_{CP} \equiv \frac{P - \overline{P}}{P + \overline{P}} \approx \frac{\Delta m_{12}^2 L}{E} \cdot \frac{\sin 2\theta_{12}}{\sin \theta_{13}} \cdot \sin \delta$$

$$N(v_e) \propto \sin^2 2\theta_{13} ; A_{CP} \propto \frac{1}{\sin \theta_{13}}$$
  
Size of  $\theta_{13}$  critical !

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#### **CPV vs matter effect**



 $@sin^{2}2\theta_{13}=0.01$ 

Smaller distance/lower energy  $\rightarrow$  small matter effect Pure CPV & Less sensitivity on sign of  $\Delta m^2$ Combination of diff. E&L help to solve.



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#### **T2K experiment** (approved & start in 2009) Long baseline neutrino oscillation experiment from Tokai to Kamioka. Sendai beam $\sim 1 \text{GeV}$ Super-K: 50 kton **J-PARC Water Cherenkov** of K2K 0.75MW 50GeV PS Super Kamiokande / 295km JAÉRI Kami (Tokai) Гokai KEK Phase2: **Phase2: Mton** Tokyo 🚽 Gifu Kanasaki Nagoya **MW** ? **Hyper-K**? Kuete Yokohama <u>Osaka</u>

## **Physics motivations**

- Discovery of  $V_{\mu \rightarrow} V_e$  appearance
- •Precise meas. of disappearance  $\nu_{\mu \rightarrow} \nu_{\textbf{X}}$
- Discovery of CP violation (Phase2)

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# Japan Proton Accelerator Research Complex (J-PARC)



# JAERI Tokai-site

February, 2004

JAERI: Japan Atomic Energy Research Institute



# **J-PARC Neutrino facility Approved in Dec. 2003** for 5 years construction (2004~2008JFY)

#### **Components**

- Primary proton beam line
  - Superconducting combined function
- Target/Horn system
- Decay volume (130m)
- **Beam dump**
- **Muon monitor**
- **Near neutrino detector (280m)**
- Second near neutrino detector (~2km): not approved yet



#### Off Axis Beam (Ref: BNL-E889 Proposal)





# **R&D** and construction of components

#### (1<sup>st</sup>) Horn inner conductor prototype





#### Superconducting combined function magnet

#### **Prototype Coil Winding**



# $\begin{array}{c} \textbf{Detector complex} \\ \hline p & \pi & \hline p & \hline p$

- Muon monitors @ ~140m
  - Fast (spill-by-spill) monitoring of beam direction/intensity

### First Front detector @280m

- Intensity, Direction,
   Spectrum, v interaction
- Second Front Detector @ ~2km (not approved)
- Far detector @ 295km
   Super-Kamiokande (50kt)



# Near Detector @280m



## $\theta_{13}$ measurement in T2K(-I) (v<sub>e</sub> appearance search)

P(vµ-ve) =  $sin^2\theta_{23} sin^22\theta_{13} sin^2(1.27 \Delta m_{23}^2L/E)$ (@ $\Delta m^2$ =2~3x10<sup>-3</sup>) signature: CC QE event (v<sub>e</sub>+n→p+e) 1ring e-like event

# BG:

beam  $v_e$  contamination (0.4% of  $v\mu$ ) mis-reconstructed  $\pi^0$  event in non-QE events ( $v_e$ +X $\rightarrow$  $v_e$ (e)+ $\pi^0$ +X')

 $sin^22\theta_{23}$ =1,  $\delta$ =0 are assumed.



# $\pi^0$ rejection



Μγγ



(OA 2.5deg, 50GeV 5yr)₂3

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# T2K 2<sup>nd</sup> phase for CPV: 4MW-PS & Hyper-K

#### 0.75MW→4MW

- Rep. rate x 2.5
  - Double RF cavities (space OK)
  - Eliminate idling time in acc. cycle
- Double # of circulating protons
  - "barrier bucket method" to avoid space charge limit
- Issues
  - Achieve first goal (0.75MW)
  - Beam loss
  - Target,...
  - (these apply also for other projects)

#### **1Mt "Hyper-Kamiokande**



2 detectors  $\times$  48m  $\times$  50m  $\times$  250m, Total mass = 1 Mton

# ~T2K-I x 100 stat

# **Sensitivity for Mixing Angle**



# **CPV measurements**

- Measure  $\nu_e$  app. for both  $\nu_\mu$  and  $\overline{\nu}_\mu$  beam
- Take asymmetry

$$A_{CP} \equiv \frac{P(\nu_{\mu} \to \nu_{e}) - P(\overline{\nu}_{\mu} \to \overline{\nu}_{e})}{P(\nu_{\mu} \to \nu_{e}) + P(\overline{\nu}_{\mu} \to \overline{\nu}_{e})} \approx \frac{\Delta m_{12}^{2}L}{E} \cdot \frac{\sin 2\theta_{12}}{\sin \theta_{13}} \cdot \sin \delta$$



#### $\overline{v}$ / v CC interaction spectrum for CPV meas.





	signal		background				
	δ <b>=0</b>	δ=π/2	total	$\nu_{\mu}$	$\nu_{\mu}$	ν <sub>e</sub>	ν <sub>e</sub>
$\nu_{\mu} \rightarrow \nu_{e}$	536	229	913	370	66	450	26
$\overline{\nu_{\mu}} \rightarrow \overline{\nu_{e}}$	536	790	1782	399	657	297	430

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 $3\sigma$  Sensitivity for CPV in T2K-II



 $3\sigma$  CP sensitivity :  $|\delta|$ >20° for sin<sup>2</sup>2 $\theta_{13}$ >0.01 with 2% syst.

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# **NO**v**A**

- Use "existing" NuMI beamline (2005~) for MINOS
- New 50kt fine grained detector @~810km and @ 12km off axis
- Liq scint. tracker & particle board absorber (1/3X<sub>0</sub>)
- 540k channel readout
- (Alternative:full active liq.sci.)
- Possible future upgrade of MI (0.4MW→2MW):Proton driver
- Proposed. (2008?~)

![](_page_29_Figure_8.jpeg)

![](_page_29_Figure_9.jpeg)

Assuming  $\Delta m^2 = 2.5 \times 10^{-3} eV^2$ Messier, v2004

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# **NOvA Physics Reach**

![](_page_30_Figure_1.jpeg)

# **Another possibility with PD**

![](_page_31_Figure_1.jpeg)

# Europe: SPL→Frejus

![](_page_32_Figure_1.jpeg)

- 4MW 2.2GeV Superconducting Proton Linac (SPL) @ CERN
- Low energy wide band (Ev~0.3GeV)
- L=130km
- Water Cherenkov 40→ 400kt (UNO)
- ~18,000 vμ CC/year/400kt θ<sub>13</sub>, CPV
- Small matter effect
- SPL in R&D, UNO in conceptual design
  - UNO Detector Conceptual Design

10%

Only optical

separation

- A Water Cherenkov Detector
- optimized for:
- Light attenuation length limit
- PMT pressure limit
- Cost (built-in staging)

#### 60x60x60m<sup>3</sup>x3 Total Vol: 650 kton Fid. Vol: 440 kton (20xSuperK) # of 20" PMTs: 56,000 # of 8" PMTs: 14,900

# **BNL-Homestake**

- 28GeV AGS upgrade to 1MW (2MW) cf current 0.1MW
- Wide band beam (0.5~6GeV)
- L=2,540km
- Mton detector
- ~13,000  $v_{\mu}$  CC/year/500kt
- Cover higher osc. maxima

Goals

 $v_e$  appearance

- Sign of  $\Delta m_{23}$
- CPV

 $\theta_{12}$ ,  $\Delta m_{12}$ 

# Possible w/ only v run at certain parameter region

• LOI written.

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![](_page_33_Figure_16.jpeg)

#### **Brookhaven to Homestake Physics Reach**

Even with only v data, CP violation and mass hierarchy are visible in some regions of parameter space.

![](_page_34_Figure_2.jpeg)

But with both v and  $\overline{v}$  running, CP precision much higher

# Reactor $\overline{v}_e$ disappearance

 $\overline{v}_{e}$  from nuclear reactor <E>~3MeV

1-P(v<sub>e</sub>  $\rightarrow$  v<sub>e</sub>) = sin<sup>2</sup>(2 $\theta_{13}$ )sin<sup>2</sup>( $\Delta m^2_{31}L/4E$ ) + O( $\Delta m^2_{21}/\Delta m^2_{31}$ ) : pure  $\theta_{13}$ 

Small systematic error (<1%) required

Identical near det @ O(100)m & far det @a few km

![](_page_35_Figure_5.jpeg)

#### **Complementarity of Reactor-Accelerator Meas.**

![](_page_36_Figure_1.jpeg)

# **Reactor Experiment Proposals**

![](_page_37_Figure_1.jpeg)

# **Double-CHOOZ**

- Twin reactor cores
  - P=2x4.2 GWth
- Two 10 tons detectors
  - 80% dodecane + 20% PXE + 0.1% Gd
  - Near: 100-200 m 60-80 mwe
  - Far: 1.05 km 300 mwe
- 3 years Sensitivity
  - 0.6% systematics
  - No signal: sin<sup>2</sup>(2θ<sub>13</sub>) < 0.02-03 (90% C.L.)
  - Signal:  $\sin^2(2\theta_{13}) > 0.04-05 (3\sigma)$
- Prospect (approved & funded in France)
  - 2007: far detector running
  - 2008: near detector running
  - Cost ~7Meuros + civil constr.

![](_page_38_Picture_15.jpeg)

![](_page_38_Picture_16.jpeg)

![](_page_39_Figure_0.jpeg)

# Summary

- Neutrino oscillation established
  - − Atm v/K2K →  $\theta_{23}$ ,  $\Delta m_{23}^2$
  - − Solar/KamLAND →  $\theta_{12}$ ,  $\Delta m_{12}^2$
- Next important issues
  - Discovery and measurement of only unknown mixing  $\theta_{13}$
  - Mass hierarchy
  - CP violation
- Future LBL exp's have good chance to achieve the goals
  - T2K using J-PARC and SK started construction. Start exp. In 2009  $\theta_{13}$  sensitivity ~0.007 (90%CL)
    - $|\delta|$ ~20deg in phase 2
  - NOvA proposal w/ similar potential to T2K
- Pure  $\theta_{13}$  measurements by reactor experiments
  - complementary to disentangle parameter relations  $\theta_{13}$  sensitivity 0.01~0.03 (90%)
  - Systematic error (<1%) is key issue
  - Double-CHOOZ is partially approved
- Neutrino will continue to be exciting for coming decades 41